Contents lists available at ScienceDirect





journal homepage: www.elsevier.com/locate/fusengdes



# Wall cratering upon high velocity normal dust impact



Panagiotis Tolias <sup>a,\*</sup>, Marco De Angeli <sup>b</sup>, Dario Ripamonti <sup>c</sup>, Svetlana Ratynskaia <sup>a</sup>, Giulio Riva <sup>c</sup>, Giambattista Daminelli <sup>c</sup>, Monica De Angeli <sup>b</sup>

<sup>a</sup> Space and Plasma Physics - KTH Royal Institute of Technology, Teknikringen 31, 10044 Stockholm, Sweden <sup>b</sup> Institute for Plasma Science and Technology, CNR, via Cozzi 53, 20125 Milano, Italy

<sup>c</sup> Institute of Condensed Matter Chemistry and Energy Technologies, CNR, via Cozzi 53, 20125 Milano, Italy

# ARTICLE INFO

Keywords: Dust in tokamaks Runaway electron termination Mechanical impacts Damage laws Impact bonding

# ABSTRACT

Dust-wall high speed impacts, triggered by the termination of runaway electrons on plasma facing components, constitute a source of erosion. Normal high velocity mechanical impacts of tungsten dust on bulk tungsten plates are reproduced in a controlled manner by light gas gun shooting systems. Post-mortem surface analysis revealed that three erosion regimes are realized; plastic deformation, bonding and partial disintegration. The large impact statistics allowed the extraction of reliable empirical damage laws in the latter regime, which can be employed for erosion estimates in future reactors.

# 1. Introduction

The presence of dust in fusion devices has important operational and safety implications, especially in future fusion reactors where metallic dust generation could scale up by orders of magnitude compared to existing devices and nuclear licensing requirements impose several restrictions on the total and hot dust inventories [1–3]. Hence, the understanding and modelling of dust transport, production, remobilization, adhesion and survivability have received wide attention within the fusion community [4–9]. Dust-wall mechanical impacts have emerged as an essential feature of dust dynamics in tokamaks [10]. As a consequence of the curved ion flow and inertial effects, mechanical impacts, that span a very wide range of incident speeds, are unavoidable. They have been revealed to control dust migration in the fusion boundary plasma [10], to influence dust survivability [11], to play a pivotal role in the formation of dust accumulation sites [12,13] and to constitute a source of plasma-facing component (PFC) damage [14].

Experimental and theoretical investigations mainly focused on the low-to-moderate impact speed range,  $v_{\rm imp} \lesssim 200 \,\mathrm{m/s}$ , being the most typical for micrometer dust-wall collisions in tokamaks [15–18]. Studies also targeted the hyper-velocity range,  $v_{\rm imp} \gtrsim 4000 \,\mathrm{m/s}$ , given the possibility for excessive wall damage, since it is characterized by complete dust vaporization, deep crater formation and fast secondary solid ejecta production [19–24]. However, hyper-velocity dust speeds are currently deemed as rather unrealistic in tokamaks. On the other hand, no investigation has concentrated on the high velocity range,  $200 \lesssim v_{\rm imp} [{\rm m/s}] \lesssim 4000$ , that is characterized by strong deformation of dust and shallow crater formation [25]. It has been recently concluded

that the high velocity range is attainable for tokamak dust [14]. Thus, high velocity impacts comprise an unexplored source of wall erosion and an unidentified mechanism of tokamak-born dust destruction.

The above naturally bring forth the question of whether high velocity tungsten dust impacts on tungsten PFCs (W-on-W) can be harmful for future fusion reactors. The first step towards an answer requires the formulation of a reliable damage law that correlates the crater volume with the dust size, impact speed and impact angle. In this work, given the lack of computational tools that account for the complex processes that unfold inside the impact site, an experimental study is performed. Nearly monodisperse W spherical dust populations are prepared that are accelerated in a controlled manner to high speeds (500-3200 m/s) towards a W plate with a light gas gun shooting system. Crater characteristics (diameter, depth) are then obtained by means of a scanning electron microscope, an optical microscope and a mechanical profiler. Finally, the large normal impact statistics are exploited for the formulation of reliable scaling laws for the crater characteristics as functions of the impact speed and dust size, which can be ultimately employed for erosion estimates in future reactors.

#### 2. Aspects of fusion relevant dust-wall impacts

Mechanical collisions between solid spherical micrometric projectiles and semi-infinite bulk solid targets have been extensively studied in the literature due to their relevance in geological phenomena, space applications, weapons research and technological applications. For refractory metal projectiles and targets (W, Mo), three impact speed

https://doi.org/10.1016/j.fusengdes.2023.113938

Received 16 January 2023; Received in revised form 24 June 2023; Accepted 27 July 2023 Available online 7 August 2023

0920-3796/© 2023 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/).

<sup>\*</sup> Corresponding author. *E-mail address:* tolias@kth.se (P. Tolias).

ranges can be distinguished that are further separated in multiple regimes [25]. In what follows, we shall present the elementary characteristics of each velocity range and briefly discuss its relevance to dust in tokamaks. It is worth pointing out that each velocity range limit depends not only on the material composition but also on the dust size and dust & wall temperature. Thus, the limits stated below should be deemed as indicative. The discussion is also applicable to oblique impacts, but not to molten projectiles or targets.

In the **low-to-moderate velocity range**,  $v_{\rm imp} \lesssim 200$  m/s, the mechanical impact induces weak plastic deformations to the projectile and the target [26]. Two regimes can be distinguished; the *sticking regime* of  $v_{\rm imp} \lesssim 5$  m/s where the irreversible adhesive work performed during the near-elastic impact dissipates the incident kinetic energy of the projectile and thus leads to immobilization [27], the *inelastic rebound regime*  $5 \lesssim v_{\rm imp} [{\rm m/s}] \lesssim 200$  where the projectile rebounds with a decreased kinetic energy due to adhesive work, plastic dissipation and frictional losses [28]. This is the most typical range for tokamak dust that has been studied experimentally and theoretically for fusion-relevant materials [15–18]. In particular, a unified analytical description has been achieved for both regimes by combining an elastic-perfectly plastic adhesive impact mechanics model [29] with a rigid sliding body model [30].

In the **high velocity range**,  $200 \leq v_{imp}[m/s] \leq 4000$ , the mechanical impact induces strong plastic deformations, surface melting and partial fragmentation to both the projectile and target [31]. Three regimes can be distinguished; the *deformation regime* of  $200 \leq v_{imp}[m/s] \leq 500$  that is accompanied by severe projectile flattening & shallow target crater formation [32], the bonding or cold spray regime of 500  $\leq v_{imp}[m/s] \leq$ 1000 that is characterized by the adhesion of the projectile on the target, which can be realized by various mechanisms such as localized melting, adiabatic shear instability, viscous interlocking and interfacial amorphization [33–37], the partial disintegration regime of 1000  $\leq$  $v_{\rm imp}[m/s] \lesssim 4000$  that is accompanied by material splash ejection and partial fragmentation [31]. It has been recently demonstrated that the high velocity range is relevant for the solid dust that is produced by the explosive stopping of runaway electrons on PFCs. In particular, postmortem and in situ evidence from FTU (2013 shutdown and 2019 decommissioning), strongly supported by dust transport simulations and by dedicated laboratory tests, revealed that the striking of runaway electrons on the equatorial poloidal limiter led to the generation of very fast solid dust that generated multiple shallow craters upon impact on the adjacent toroidal limiters [14].

In the hyper-velocity range of  $v_{\rm imp} \gtrsim 4000$  m/s, the impact velocity is comparable or larger than the compressive sound speeds of the projectile and the target [25]. This triggers the emergence of strong shockwaves, upon whose sudden compaction extreme transient pressures and temperatures arise at the collision zone. Upon shockwave release, both materials extensively vaporize while the localized energy is large enough to cause atomic excitation and partial ionization [38]. As a result, hyper-velocity impacts are accompanied by strong electrostatic [38,39] and spectroscopic signals [40,41]. In addition, the emerging craters feature excavated volumes that by far exceed the volume of the projectile. Finally, very high velocity solid ejecta can be released that cause further damage to the surrounding material [42]. Solid evidence of hypervelocity impacts have been obtained in FTU that concerned the in-situ detection of dust impact ionization by electrostatic probes and the post-mortem observation of craters [19-21]. However, considering the typical initial dust velocity distributions and given the generally restricted acceleration lengths, no acceleration mechanisms are currently known that can lead to impact speeds within the hyper-velocity range [4,8].

# 3. Experimental

High-sphericity low internal porosity W dust was purchased from "TEKNA Plasma Systems". The original batch had a nominal size



**Fig. 1.** Image of a damaged bulk W target after the high velocity impact of spherical W dust (here  $D_d = 51 \,\mu\text{m}$  and  $v_{imp} = 2075 \,\text{m/s}$ ) together with magnified SEM images of two regions that have been included in the statistical crater analysis.

# Table 1

W-on-W impacts realized by the one- and two-stage light gas gun. In what follows,  $D_{\rm d}$  denotes the mean dust diameter,  $v_{\rm imp}$  the impact speed, # the number of craters,  $D_{\rm c}$  the measured crater diameter,  $\sigma_{\rm D}$  the standard deviation in the crater diameter,  $H_{\rm c}$  the measured crater depth,  $\sigma_{\rm H}$  the standard deviation in the crater depth. The negative sign in the crater depth is indicative of impact bonding.

Dd	$v_{\rm imp}$	#	D <sub>c</sub>	$\sigma_{ m D}$	$H_{\rm c}$	$\sigma_{ m H}$
(µm)	(m/s)		(µm)	(µm)	(µm)	(µm)
51	660	97	74.2	15.5	+7.5	3.3
51	984	83	73.0	15.3	-13.3	3.9
51	1565	62	84.0	17.7	+14.9	5.7
51	2075	73	99.3	20.9	+32.2	5.1
51	2506	47	106.1	22.3	+40.4	5.3
51	2561	36	107.8	22.9	+38.4	7.1
51	3128	37	115.7	24.4	+50.0	6.7
63	596	110	84.2	17.5	+8.1	3.2
63	764	64	89.0	18.4	-26.3	4.2
63	1012	63	89.2	18.2	-14.3	4.6
63	1534	72	97.8	20.5	+21.4	6.4
63	2039	78	119.0	24.8	+43.3	7.3
63	2485	37	129.3	26.8	+51.1	8.1
63	2500	45	130.9	27.4	+47.5	8.2
63	3108	33	142.0	35.1	+62.4	11.3
63	3190	36	149.3	30.8	+66.6	8.2
76	583	99	104.9	21.3	+8.7	3.2
76	997	48	108.2	22.2	-16.8	5.8
76	1563	57	121.8	26.6	+31.6	10.5
76	2033	35	143.3	29.4	+50.1	7.3
76	2058	43	145.6	30.7	+48.1	5.8
76	2513	25	153.8	31.4	+64.4	8.9
76	2551	36	152.6	33.4	+65.3	10.7
76	3069	27	180.7	42.3	+81.9	14.2
76	3126	20	180.0	38.4	+80.1	9.2

distribution of  $45 - 90 \,\mu\text{m}$ . From this polydisperse batch, three nearly monodisperse sub-populations were meshed out using a sequence of six sieves with nominal sizes of 80, 75, 71, 63, 56 and 50  $\mu$ m. The mean W dust diameters are 51 ( $\pm$ 5)  $\mu$ m, 63 ( $\pm$ 5)  $\mu$ m and 76 ( $\pm$ 5)  $\mu$ m, sizes that are comparable to the most probable sizes of the fast Mo solid dust that was observed in FTU [14]. The bulk square W targets have 23 mm length and 4 mm thickness.

High velocity dust-wall impacts are realized by means of a light gas gun system [43,44]. In the *two stage configuration*, the first stage is a high pressure reservoir connected to the second stage with a fast valve. The second stage (or pump tube) is a cylinder in which the light



**Fig. 2.** SEM images of the three erosion regimes that are realized in high velocity W-on-W normal dust-wall impacts for a given spherical dust size ( $D_d = 63 \,\mu\text{m}$ ): (a) **plastic deformation regime** (here for  $v_{imp} = 596 \,\text{m/s}$ ), tilted image at 45°, (b) **impact bonding regime** (here for  $v_{imp} = 764 \,\text{m/s}$ ), tilted image at 45°, (c) **partial disintegration regime** (here for  $v_{imp} = 1563 \,\text{m/s}$ ), normal image at 90°. Note the subtle morphological differences between (a) and (c); in the partial disintegration regime, the crater rim is rougher as well as more elevated and small fragments or melt splashes are present within the crater cone.

gas is fed at relatively low pressure and compressed by a free piston following the fast valve operation. The compressed light gas rapidly expands into the launch tube, which has a smaller diameter than the pump tube, simultaneously accelerating a macro-scale projectile (sabot) which features a cavity that is loaded with micron dust. The dust particles are separated from the sabot at the end of the launch tube where the latter impinges on a diaphragm. The dust particles free stream with high speeds into a vacuum chamber that features the target. Impact speeds within the range of  $1000 \leq v_{imp}[m/s] \leq 3500$  have been realized by employing different light gases (nitrogen or hydrogen) and setting different initial pressures for the first (20–40 bar) & second (0.8–1.2 bar) stage. In the *single stage configuration*, the high-pressure hydrogen reservoir is directly connected to the launch tube without an intervening piston and the achieved impact speeds lie within  $400 \leq v_{imp}[m/s] \leq 1000$ .

The dust speed is measured through the dust transit time between two laser sheets with an uncertainty that is generally less than  $\pm 1\%$ . This measurement error primarily originates from the relatively large thickness of the individual beams (1 mm) when compared to the beam spacing (100 mm). The optical method was preferred over the timeof-flight method [44,45], where the start trigger is provided by the muzzle flash due to sabot ejection from the barrel and the stop trigger is provided by the impact flash, whose precision is  $\sim \pm 3\%$ . The dust cloud speed should well approximate the dust particle speed, given the near-constant small cloud width recorded by the laser sheets.

Overall, 34 impact tests were performed with the three dust subpopulations and impact speeds covering the high velocity range. Light gas gun acceleration leads to a large number of W-on-W impacts, since it is not possible to load a small dust number on the sabot cavity. However, the useful crater statistics are rather restricted, since the central spot features many overlapping craters that should be excluded from the analysis (see Fig. 1) and since the central spot also features few visibly smaller craters due to dust-dust impacts in the proximity of the target (see the finite width of the dust cloud). Naturally, it is possible to minimize the central spot by loading less dust on the sabot, but this compromises the accuracy of the speed measurement and does not necessarily improve statistics. From the 34 shot samples, 25 featured a large enough number of isolated craters suitable for statistical analysis (Table 1).

The shot W-on-W samples were first mapped by means of a Scanning Electron Microscope (SEM), at low magnification and at high resolution, in order to statistically estimate the average value of the crater diameter at different impact conditions. The crater depth was then measured with a precision optical microscope (Leitz Wetzlar Ortolux) of 0.5  $\mu$ m sensitivity. Moreover, few craters from each shot sample were mapped with a mechanical profiler (KLA-Tencor mod. P-15) in order to verify the precision of the above techniques. The instrumental uncertainty was estimated to be  $\pm 3 \,\mu$ m for the crater depth (optical) and  $\pm 15\%$  for the crater diameter (SEM). The average values and standard deviations of the crater dimensions (instrumental and statistical errors) are listed in Table 1.

Since the impact speeds varied from 583 m/s to 3190 m/s, all three impact regimes of the high velocity range were observed for all the three dust sub-populations. Characteristic examples are illustrated in Fig. 2.

#### 4. Analysis

# 4.1. Empirical damage laws

Impact damage laws are empirical correlations of the crater diameter and depth as function of the target and projectile material properties, impact speed, impact angle and dust size [46–48]. Systematic hyper-velocity dust impact experiments have led to the formulation of empirical damage laws that are applicable in quite extended impact velocity and dust size ranges [46–52]. Considering their validity for various target-projectile combinations, such general damage laws are of limited accuracy, especially in the high velocity impact range. In what follows, we utilize our extensive W-on-W experimental data to extract empirical damage laws valid within the disintegration regime of the high velocity range,  $1 \leq v_{imp} [km/s] \leq 3.5$ .

The experimental crater depth and crater diameter data are fitted in the standard power law form of  $\alpha (D_d)^\beta (v_{imp})^\gamma$  with  $\alpha, \beta, \gamma$  the fitting parameters. The two-variable three-parameter non-linear fit led to the empirical damage laws

$$D_{\rm c} = 0.0330 (D_{\rm d})^{1.005} (v_{\rm imp})^{0.527}, \tag{1}$$

$$H_c = 0.0000114(D_d)^{1.264}(v_{\rm imp})^{1.282},$$
(2)

with  $D_c$ ,  $H_c$ ,  $D_d$  measured in  $\mu$ m and  $v_{imp}$  in m/s. The mean absolute relative fitting error is 1.87% for the diameter and 6.89% for the depth. The standard errors for the fitting parameters of the diameter are 0.0089 ( $\alpha$ ), 0.041 ( $\beta$ ), 0.027 ( $\gamma$ ). The standard errors for the fitting parameters of the depth are 0.00000803 ( $\alpha$ ), 0.098 ( $\beta$ ) and 0.072 ( $\gamma$ ). The empirical damage laws are plotted in Fig. 3.

An established general damage law for the crater depth, appropriate for metallic dust exceeding 50 µm and impact speeds within 2  $\lesssim v[\text{km/s}] \lesssim 12$ , reads as [46–48]

$$H_{\rm c} = 5.24 \frac{D_{\rm d}^{19/18}}{h_{\rm t}^{1/4}} \left(\frac{\rho_{\rm d}}{\rho_{\rm t}}\right)^{1/2} \left(\frac{v_{\rm imp}}{c_{\rm t}}\right)^{2/3},\tag{3}$$

where  $H_c$  is the crater depth in cm,  $D_d$  the dust diameter in cm,  $h_t$  the target's Brinell hardness number,  $\rho_d$ ,  $\rho_t$  the mass densities of the dust and the target,  $c_t$  the target's sound speed and  $v_{imp}$  the dust impact speed. It is evident from the exponents that the general damage law exhibits a weaker dependence on both the dust size and the impact speed than the W-on-W damage law. More specifically, as discerned from Fig. 4, the general damage law significantly underestimates the crater depth for large dust sizes and high impact speeds. This confirms the importance of target-projectile specific damage laws as those of Eq. (1),(2).



Fig. 3. The crater diameter (top panel) and crater depth (bottom panel) within the *partial disintegration regime* as functions of the normal W-on-W speed for three spherical W dust sub-populations. Experimental results (purple symbols with error bars) together with the predictions of the empirical damage laws of Eq. (1),(2) (solid black lines).



**Fig. 4.** The crater depth in the *partial disintegration regime* of the high velocity range as a function of the W-on-W impact speed for two dust sizes. Comparison of the predictions of our specific damage law, Eq. (2), with the general damage law, Eq. (3).

Finally, empirical damage laws allow for an estimate of the excavated volume provided that an additional assumption is made for the geometry of the crater. As expected from symmetry considerations arising from the high dust sphericity, local target planarity and normal nature of the impact, the crater geometry can be well-approximated by a spherical cap. This has been consistently observed in the laboratory impact tests, see Fig. 5, and is consistent with the experimental hyper-velocity literature [53,54]. Thus, the excavated volume can be approximated by

$$V_{\rm c} = \frac{1}{6}\pi H_{\rm c} \left(\frac{3}{4}D_{\rm c}^2 + H_{\rm c}^2\right).$$
<sup>(4)</sup>

Some characteristic examples are provided in Fig. 6. Note that the excavated material from a single impact crater is much larger than the excavated material from a single unipolar arc crater [55,56], but much smaller than the material excavated during an edge-localized mode or major disruption driven melt event [57,58].

# 4.2. Critical velocities

Here, we shall discuss the transition between the high velocity regimes for normal W-on-W impacts, which dictates the applicability limits of our empirical damage laws.

The critical erosion velocity  $v_{\text{crit}}^{\text{ero}}$  is the transition velocity between the bonding and disintegration regimes of the high velocity range. An analytical expression has been derived for  $v_{\text{crit}}^{\text{ero}}$  that associates the transition with the impact kinetic energy that triggers melting [31]. A first-order

energy balance analysis that contains a number of approximations leads to the expression [31]

$$p_{\rm crit}^{\rm ero} = \left(\frac{20e_{\rm th}I_{\rm melt}}{\rho_{\rm d}\sqrt{D_{\rm d}}}\right)^{2/5}$$

where  $e_{\rm th} = \sqrt{\rho_{\rm d}k_{\rm d}c_{\rm pd}}$  is the so-called thermal effusivity with  $k_{\rm d}$  the thermal conductivity and  $c_{\rm pd}$  the specific isobaric heat capacity, where  $I_{\rm melt} = T_{\rm m} - T_0 + \Delta h_{\rm f}/c_{\rm pd}$  is a melting index with  $T_{\rm m} = 3695$  K the melting point of W,  $T_0 = 300$  K the room temperature and  $\Delta h_{\rm f}$  the latent heat of fusion. This expression should be expected to underestimate the critical erosion velocity but has exhibited a good agreement with dedicated Sn, Bi, Zn, Ti experiments [31].

The *critical bonding velocity*  $v_{\text{crit}}^{\text{bond}}$  is the transition velocity between the deformation and bonding regimes of the high velocity range. A closed-form correlation is available for  $v_{\text{crit}}^{\text{bond}}$  that is based on numerical simulations within the assumption that impact bonding can be attributed solely to the onset of adiabatic shear instabilities [33]. Additional numerical modelling and experimental results allowed the consideration of size effects in the original expression [34]. The final expression reads as [33,34]

$$v_{\text{crit}}^{\text{bond}} = \left[ 667 - 0.014 \rho_{\text{d}} + 0.08(T_{\text{m}} - T_{0}) + 10^{-7} \sigma_{\text{y}} \right] (D_{\text{d}})^{-0.07}$$

where all quantities are in SI units and  $\sigma_y$  denotes the yield strength. It is worth pointing out that the adopted size exponent is a factor of two less than its traditional value [34] in accordance with state-of-the-art experiments [35].

In Fig. 7, the two W-on-W critical velocities are plotted as a function of the dust size. The room temperature recommendations of Ref. [59] have been followed for the W thermophysical properties. Velocity-size combinations that have been experimentally verified to lead to impacts that belong to the deformation and bonding regimes have also been included. It is evident that the aforementioned general expressions underestimate both critical velocities.

# 5. Summary and future work

The first experimental study of high velocity W dust - W wall impacts, relevant for cascade-like PFC damage caused by runaway electron stopping on PFCs, has been carried out. Controlled normal high velocity impacts have been realized with a light gas dust gun. Surface analysis allowed the identification of three wall damage regimes (plastic deformation, impact bonding, partial disintegration) and the extraction of empirical damage laws for the most harmful disintegration regime. Work in progress focuses on the extension of the impact data



**Fig. 5.** SEM images of high velocity W-on-W normal dust impacts that fall into the *partial disintegration regime* for different dust sizes and different impact speeds: (a)  $D_d = 51 \,\mu\text{m}, v_{imp} = 1506 \,\text{m/s}$ , (b)  $D_d = 63 \,\mu\text{m}, v_{imp} = 3190 \,\text{m/s}$ , (c)  $D_d = 76 \,\mu\text{m}, v_{imp} = 2033 \,\text{m/s}$ .



**Fig. 6.** The excavated volume predictions of the empirical damage laws for the crater diameter and the crater depth in the *partial disintegration regime* under the spherical cap assumption. Excavated volume as a function of the impact speed for various W dust sizes.



Fig. 7. Critical erosion (purple line) and bonding velocities (blue line) versus the dust size for normal W-on-W impacts. Impacts within the deformation (green stars) and bonding regime (red circles).

to obtain more accurate damage laws, the determination of the critical velocities that demarcate the damage regimes and the realization of oblique impacts. In the future, we aim to quantify the effect of elevated dust temperatures; a step that is particularly important given that explosive runaway electron termination on PFCs is expected to produce fast hot solid dust.

Concerning the potential significance of high velocity solid dust impacts in fusion devices, the following remark is important. While primary runaway electron induced damage could have a major extent, it has a localized character and it could be constrained to sacrificial limiters or to replaceable divertor plates in future fusion reactors. On the other hand, secondary high velocity dust impact damage should be considerably less dramatic, but it has a delocalized character and could thus compromise the integrity of plasma facing components including diagnostics.

# Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

# Data availability

Data will be made available on request.

# Acknowledgements

The work has been performed within the framework of the EU-ROfusion Consortium, funded by the European Union via the Euratom Research and Training Programme (Grant Agreement No101052200 - EUROfusion). Views and opinions expressed are however those of the authors only and do not necessarily reflect those of the European Union or European Commission. Neither the European Union nor the European Commission can be held responsible for them.

# References

- G. Federici, C.H. Skinner, J.N. Brooks, J.P. Coad, et al., Nucl. Fusion 41 (2001) 1967.
- [2] J.P. Sharpe, D.A. Petti, H.-W. Bartels, Fusion Eng. Des. 63-64 (2002) 153.
- [3] J. Roth, E. Tsitrone, A. Loarte, Th. Loarer, et al., J. Nucl. Mater. 390-391 (2009) 1.
- [4] S.I. Krasheninnikov, R.D. Smirnov, D.L. Rudakov, Plasma Phys. Control. Fusion 53 (2011) 083001.
- [5] P. Tolias, S. Ratynskaia, M. De Angeli, G. De Temmerman, et al., Plasma Phys. Control. Fusion 58 (2016) 025009.
- [6] J.T. Holgate, L. Simons, Y. Andrew, C.K. Stavrou, Europhys. Lett. 127 (2019) 45004.
- [7] M. De Angeli, E. Lazzaro, P. Tolias, S. Ratynskaia, et al., Nucl. Fusion 59 (2019) 106033.

#### P. Tolias et al.

- [8] S. Ratynskaia, L. Vignitchouk, P. Tolias, Plasma Phys. Control. Fusion 64 (2022) 044004.
- [9] S. Ratynskaia, A. Bortolon, S.I. Krasheninnikov, Rev. Mod. Plasma Phys. 6 (2022) 20.
- [10] S. Ratynskaia, L. Vignitchouk, P. Tolias, I. Bykov, et al., Nucl. Fusion 53 (2013) 123002.
- [11] A. Shalpegin, L. Vignitchouk, I. Erofeev, F. Brochard, et al., Plasma Phys. Control. Fusion 57 (2015) 125017.
- [12] L. Vignitchouk, K. Paschalidis, S. Ratynskaia, P. Tolias, R.A. Pitts, Plasma Phys. Control. Fusion 65 (2023) 015014.
- [13] L. Vignitchouk, S. Ratynskaia, P. Tolias, R.A. Pitts, et al., Nucl. Fusion 58 (2018) 076008.
- [14] M. De Angeli, P. Tolias, S. Ratynskaia, D. Ripamonti, et al., Nucl. Fusion 63 (2023) 014001.
- [15] L. Vignitchouk, P. Tolias, S. Ratynskaia, Plasma Phys. Control. Fusion 56 (2014) 095005.
- [16] S. Ratynskaia, P. Tolias, A. Shalpegin, L. Vignitchouk, et al., J. Nucl. Mater. 463 (2015) 877.
- [17] A. Shalpegin, F. Brochard, S. Ratynskaia, P. Tolias, et al., Nucl. Fusion 55 (2015) 112001.
- [18] P. Tolias, S. Ratynskaia, A. Shalpegin, L. Vignitchouk, et al., Nucl. Mater. Energy 12 (2017) 524.
- [19] C. Castaldo, S. Ratynskaia, V. Pericoli, U. de Angelis, et al., Nucl. Fusion 47 (2007) L5.
- [20] S. Ratynskaia, C. Castaldo, K. Rypdal, G. Morfill, et al., Nucl. Fusion 48 (2008) 015006.
- [21] S. Ratynskaia, C. Castaldo, E. Giovannozzi, D. Rudakov, et al., Plasma Phys. Control. Fusion 50 (2008) 124046.
- [22] R.D. Smirnov, S.I. Krasheninnikov, A. Yu. Pigarov, D.J. Benson, et al., J. Nucl. Mater. 390-391 (2009) 84.
- [23] A. Fraile, P. Dwivedi, G. Bonny, T. Polcar, Nucl. Fusion 62 (2022) 026034.
- [24] G.-J. Niu, Q. Xu, Z.-S. Yang, T. He, et al., J. Nucl. Mater. 572 (2022) 154048.
- [25] S.V. Klinkov, V.F. Kosarev, M. Rein, Aerosp. Sci. Technol. 9 (2005) 582.
- [26] W. Stronge, Impact Mechanics, Cambridge University Press, Cambridge, 2000.
- [27] W. John, Aerosol Sci. Technol. 23 (1995) 2.
- [28] R.M. Brach, P.F. Dunn, X. Li, J. Adhes. 74 (2000) 227.
- [29] C. Thornton, Z. Ning, Powder Technol. 99 (1998) 154.
- [30] D. Gorham, A. Kharaz, Powder Technol. 112 (2000) 193.
- [31] M. Hassani-Gangaraj, D. Veysset, K.A. Nelson, C.A. Schuh, Nature Commun. 9 (2018) 5077.
- [32] M. Hassani-Gangaraj, D. Veysset, K.A. Nelson, C.A. Schuh, Phys. Rev. Lett. 119 (2017) 175701.

- [33] H. Assadi, F. Gärtner, T. Stoltenhoff, H. Kreye, Acta Mater. 51 (2003) 4379.
- [34] T. Schmidt, F. Gärtner, H. Assadi, H. Kreye, Acta Mater. 54 (2006) 729.
- [35] M. Hassani-Gangaraj, D. Veysset, K.A. Nelson, C.A. Schuh, Scr. Mater. 145 (2018) 9.
- [36] M. Hassani-Gangaraj, D. Veysset, K.A. Nelson, C.A. Schuh, Appl. Surf. Sci. 476 (2019) 528.
- [37] Y. Sun, D. Veysset, K.A. Nelson, C.A. Schuh, Appl. Phys. Lett. 117 (2020) 134105.
   [38] M.J. Burchell, M.J. Cole, J.A.M. McDonnell, J.C. Zarnecki, Meas. Sci. Technol.
- 10 (1999) 41.
- [39] N. Lee, S. Close, D. Lauben, I. Linscott, et al., Int. J. Impact Eng. 44 (2012) 40.
  [40] G. Eichhorn, Planet. Space Sci. 24 (1976) 771.
- [41] D. Heunoske, M. Schimmerohn, J. Osterholz, F. Schäfer, Procedia Eng. 58 (2013)
   624
- [42] G. Eichhorn, Planet. Space Sci. 26 (1978) 469.
- [43] G. Riva, A. Reggiori, Fusion Technol. 15 (1989) 143.
- [44] D. Veysset, J.-H. Lee, M. Hassani, S. Kooi, E. Thomas, K. Nelson, Appl. Phys. Rev. 8 (2021) 011319.
- [45] F.K. Schäfer, T. Geyer, E.E. Schneider, M. Rott, E. Igenbergs, Int. J. Impact Eng. 26 (2001) 683.
- [46] N. Pailer, E. Grün, Planet. Space Sci. 28 (1980) 321.
- [47] B.G. Cour-Palais, Int. J. Impact Eng. 5 (1987) 221.
- [48] L. Berthoud, J.C. Mandeville, Proceedings of the First European Conference on Space Debris, 1993, p. 459.
- [49] M. Lambert, Adv. Space Res. 19 (1997) 369.
- [50] K.R. Housen, K.A. Holsapple, Icarus 211 (2011) 856.
- [51] F. Hörz, Meteorit. Planet. Sci. 47 (2012) 763.
- [52] A.J. Gemer, Z. Sternovsky, D. James, M. Horanyi, Planet. Space Sci. 183 (2020) 104628.
- [53] D.E. Gault, The Moon 6 (1973) 32.
- [54] J.R. Baker, Int. J. Impact Eng. 17 (1995) 25.
- [55] H.T.C. Kaufmann, C. Silva, M.S. Benilov, Plasma Phys. Control. Fusion 61 (2019) 095001.
- [56] V. Rohde, M. Balden, R, Neu and the ASDEX upgrade team, Nucl. Mater. Energy 29 (2021) 101083.
- [57] S. Ratynskaia, E. Thorén, P. Tolias, R.A. Pitts, et al., Nucl. Fusion 60 (2020) 104001.
- [58] E. Thorén, S. Ratynskaia, P. Tolias, R.A. Pitts, et al., Plasma Phys. Control. Fusion 63 (2021) 035021.
- [59] P. Tolias, Nucl. Mater. Energy 13 (2017) 42.