

1 ***Interplay and feedback between tectonic regime, faulting, sealing horizons, and fluid flow in a***  
2 ***hydrocarbon-hosting extensional basin: the Val d'Agri Basin case, southern Italy***

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13 **Keywords:**

14 extensional basin; hydrocarbon reservoir; seal integrity; carbonate clumped isotopes; active faulting; fluid  
15 circulation systems.

16

17 **Abstract**

18 Understanding the factors that govern past fluid circulation in tectonically active and/or hydrocarbon-rich  
19 basins is crucial for elucidating present-day fluid-flow scenarios. We investigate the circulation of paleo-fluids  
20 in the extensional-transtensional Val d'Agri Basin (southern Italy), home to a giant oil field and significantly  
21 affected by both natural and human-induced seismicity. Our aim is to understand how faulting and the  
22 variable thickness of the clay-rich tectonic mélange, which constitutes the seal of the hydrocarbon reservoir,  
23 influenced past fluid flow under different tectonic regimes. To achieve this, we combined multiscale  
24 structural observations with isotope (C, O, clumped, and Sr) and Rare Earth and Yttrium (REY) analyses of  
25 fault-related calcite mineralizations. Using analytical methodologies that allow the analysis of sub-milligram  
26 samples for carbonate clumped isotopes, we provided a detailed characterization of the variability in  
27 precipitation temperatures and composition of parental fluids in both space and time. Our results reveal five

28 main types of parental fluids, ranging from meteoric to intraformational and deep crustal, which were  
29 differently involved in the tectonic evolution of the Val d'Agri Basin. During orogenic shortening, vertical fluid  
30 circulation was mostly limited and compartmentalized, whereas post-orogenic extensional faulting promoted  
31 the ascent of deep fluids. Our findings indicate that the sealing properties of the *mélange* were likely  
32 enhanced locally by increased thickness but were also compromised by fault activity and associated seismic  
33 events. Fluid circulation in the study area has been influenced by the prevailing tectonic regime (compressive  
34 vs. extensional), stratigraphic-structural architecture, properties of impermeable horizons, and seismic  
35 events. The model proposed for the Val d'Agri Basin elucidates past processes that are useful for  
36 understanding current fluid circulation in the basin itself and can be applied to other basins where fluid  
37 circulation is partly manipulated by human activities.

## 38 **1. Introduction**

39 In various tectonic settings (e.g., extensional basins, fold-and-thrust belts, etc.), accumulation of fluids in the  
40 subsurface, including groundwater, hydrocarbons, and CO<sub>2</sub> is ruled by stratigraphical, lithological and  
41 mechanical properties of the sealing horizons (Grunau, 1987). These impermeable horizons, which are often  
42 clay-rich, maintain fluid confinement potentially causing fluid overpressures (Hager et al., 2021). The sealing  
43 capacity of impermeable horizons can be compromised by fault zones, which can act as fluid conduits (Caine  
44 et al., 1996; Faulkner et al., 2010) or cause fault-valve actions (Sibson, 2000; Doglioni et al., 2014). In  
45 tectonically active regions, changes in pore fluid pressure and its diffusion can trigger both natural (Miller et  
46 al., 2004) and human-induced (Ellsworth, 2013) earthquakes. Consequently, understanding the complex  
47 interaction between fluids, faults, and impermeable horizons is of significant economic and social  
48 importance, particularly for fluid exploitation (Song et al., 2022), storage (Shukla et al., 2010), and assessment  
49 of seismic and environmental hazards (Chiodini et al., 2020).

50 The Val d'Agri Basin in southern Italy serves as an outstanding natural laboratory for investigating the above-  
51 mentioned geological processes. It is a seismically active intermontane basin formed since the Middle-Late  
52 Pleistocene due to post-orogenic extensional-transtensional faulting, which dissected and partly reactivated  
53 pre-existing tectonic structures (Patacca & Scandone, 2007). The structural architecture of this basin,

54 resulting from its polyphasic tectonic history, led to the genesis of the largest onshore oil field in western  
55 Europe. Hydrocarbons are trapped below a clay-rich, fluid-overpressured, and impermeable sealing horizon,  
56 known as the Irpinia mélange (Mazzoli et al., 2001; Hager et al., 2021). Both lateral distribution and vertical  
57 integrity of the Irpinia mélange are subjects of an extensive debate (Brozzetti, 2011; Candela et al., 2015;  
58 D'Adda et al., 2017). Additionally, the Val d'Agri Basin has experienced historical strong earthquakes ( $M \leq 7$ ;  
59 Cucci et al., 2004) and recent low-magnitude seismicity ( $M \leq 3$ ; Improta et al., 2017) of both natural and  
60 induced origins.

61 In this study, we elucidate the interplay and feedback between fluid flow, faulting, and the Irpinia mélange.  
62 We focus on the evolution of paleo-fluid circulation both in space and time to deepen our understanding of  
63 current fluid circulation and envision future scenarios. We carry out a structural-geochemical study of syn-  
64 tectonic calcite mineralizations collected within fault zones of the Val d'Agri Basin to reconstruct a basin-  
65 scale, spatiotemporal conceptual model of fluid flow. Through geochemical analyses, including the largest  
66 database of carbonate clumped isotopes applied to a single case study so far, we reveal significant variability  
67 in fault-fluid interaction processes. These insights are valuable for evaluating the current integrity of sealing  
68 horizon in the Val d'Agri area, and in other hydrocarbon-rich post-orogenic extensional basins worldwide  
69 (e.g., Pannonian Basin, Hungary, Czauer & Madl-Szonyi, 2011; Bohai Bay Basin, China, Song et al., 2022;  
70 Nispiro Field, Mexico, Bourdet et al., 2010).

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## 73        2. Geological Setting

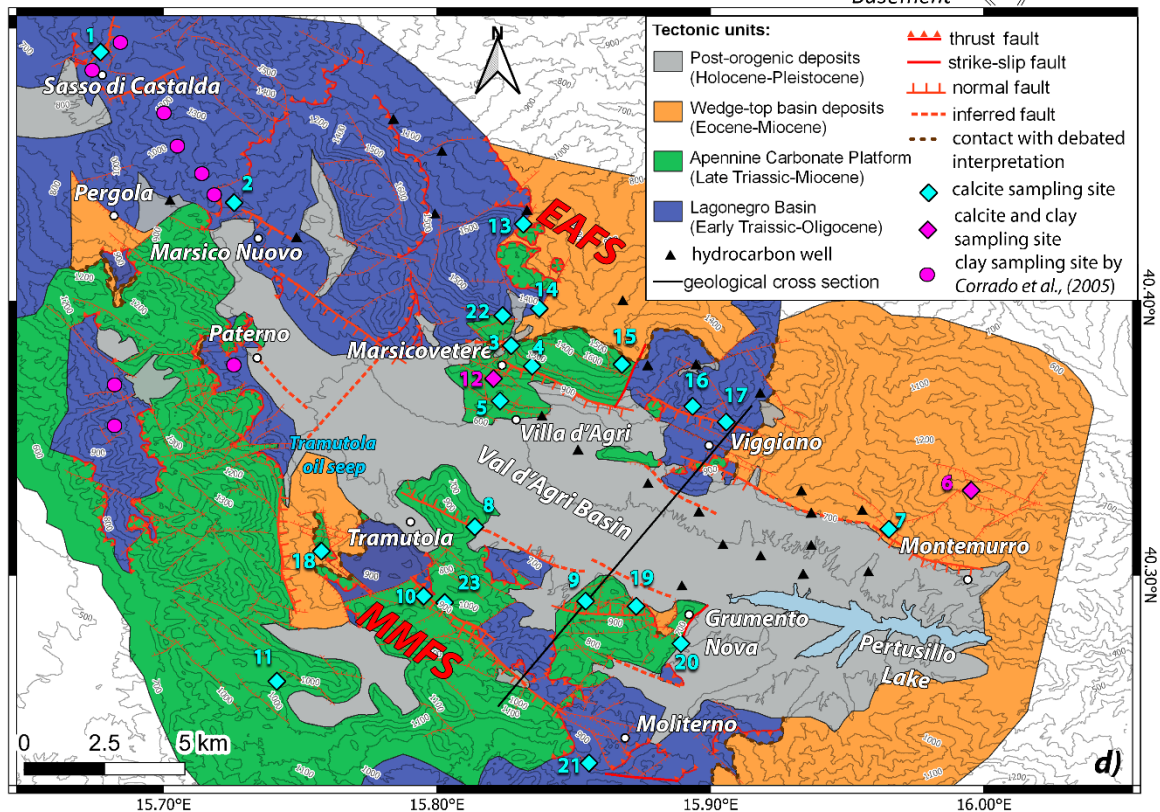
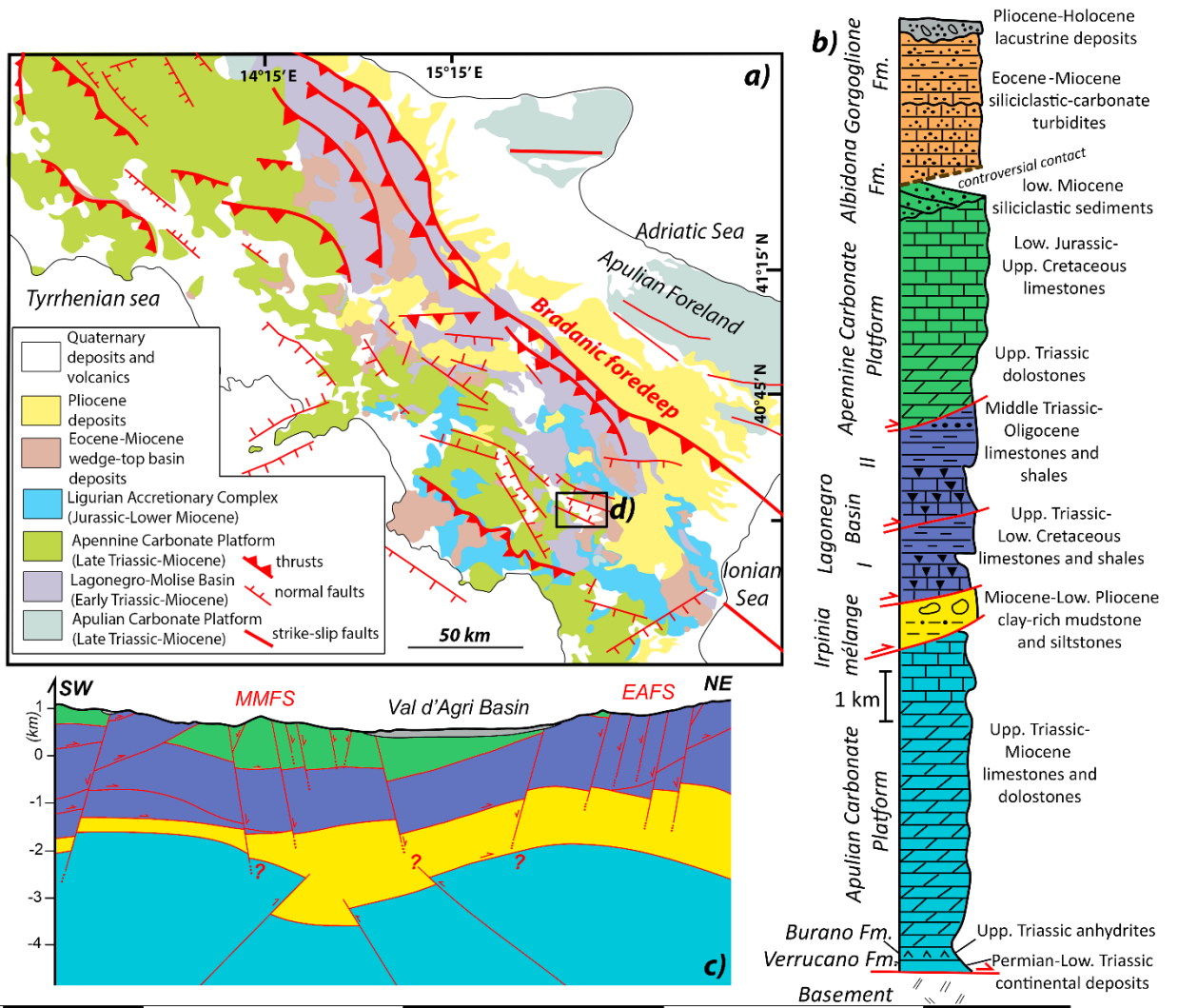
74    The Val d'Agri is a post-orogenic extensional-transtensional basin located in the Southern Apennines fold-  
75    and-thrust belt, Italy. The Southern Apennines ensued from the superimposition of orogenic and post-  
76    orogenic tectonic phases, which were associated with the eastward retreat of the westward-subducting  
77    Adriatic slab, and with the opening of the Tyrrhenian back-arc basin (e.g., Malinverno and Ryan, 1986). The  
78    orogenic shortening resulted in the stacking of tectonic units derived from the Mesozoic oceanic (Ligurian  
79    Accretionary Complex) and Adriatic passive margin domains (i.e., Apennine and Apulian Carbonate Platforms,  
80    Lagonegro Basin; Patacca and Scandone, 2007; Fig. 1), which were differentiated during Mesozoic rift-related  
81    extensional tectonics and subsequent oceanization (Patacca and Scandone, 2007). Since the Miocene time,  
82    these units were detached from their original substratum, and overthrust toward the foreland with a ca.  
83    NE-directed motion. The final orogenic phase (Late Pliocene-Early Pleistocene) was associated with a shift  
84    from thin- to thick-skinned tectonics (Shiner et al., 2004; Butler et al., 2004). Middle Pleistocene-Holocene,  
85    post-orogenic, extensional-transtensional tectonics led to the displacement or reactivation of pre-existing  
86    structures and gave rise to numerous intermontane basins including the Val d'Agri Basin (Patacca and  
87    Scandone, 2007). Within the Val d'Agri basin area, the superimposition of distinct thrust sheets, characterized  
88    by varying lithology, thickness, rheology, and permeability (Fig. 1; Patacca and Scandone, 2007) determined  
89    the genesis of the largest onshore oil field in western Europe. Its tectono-stratigraphic succession includes,  
90    from top to bottom: siliciclastic-carbonate turbidites (Albidona and Gorgoglione fms.), shallow-marine  
91    carbonates of the Apennine Platform, proximal-to-distal pelagic successions of the Lagonegro Basin, the  
92    Irpinia mélange, and shallow-marine carbonates of the Apulian Platform (Shiner et al., 2004; Patacca &  
93    Scandone, 2007; Palladino et al., 2023; Fig. 1b). The Irpinia mélange, a heterogeneous, tectonic, clay-rich, 0.1  
94    to~ 1 km thick (Mazzoli et al., 2001) horizon with a general permeability of  $< 10^{-7}$  mD (Hager et al., 2021),  
95    forms the seal of the Val d'Agri hydrocarbon reservoir constituted by the Apulian Carbonate Platform.  
96    The basin is bordered by two main fault systems: the East Agri Fault System (EAFS; Fig. 1d), to the northeast,  
97    and the Monti della Maddalena Fault System (MMFS; Fig.1d), to the southwest. Both fault systems are  
98    primarily composed of steeply dipping, NW-SE and NE-SW striking, extensional-transtensional faults (e.g.,  
99    Giano et al., 2000; Schirripa Spagnolo et al., 2024). These structures are consistent with the present NE-SW

100 oriented tectonic extension (Mariucci and Montone, 2020). The two fault systems are secondarily composed  
101 of faults with a wide range of orientations. While some authors consider all these faults as the result of the  
102 polyphase tectonic history (e.g., Cello et al., 2000), a recent model proposes a polygonal-like style of faulting  
103 controlled by the Irpinia mélangé (Schirripa Spagnolo et al., 2024). However, the role of this mélangé during  
104 post-orogenic extension is not well understood. Some studies contend that the Val d'Agri faults root along  
105 the Irpinia mélangé, decoupling the Apulian Carbonate Platform from the overlying units (D'Adda et al., 2017;  
106 Hager et al., 2021). This resulted in the soft-linkage between shallow extensional faults and deep inherited  
107 structures within the Apulian Carbonate Platform (Borraccini et al., 2002; Candela et al., 2015). Conversely,  
108 others suggest that the main faults of both EAFS and MMFS cut across and displace the Irpinia mélangé  
109 (Brozzetti, 2011; Valoroso et al., 2023).

110 Schirripa Spagnolo et al. (2024) performed the first radiometric dating of the Val d'Agri fault activity, yielding  
111 two reliable U–Pb ages. In particular, they dated slip along two extensional faults. One slickenfiber was dated  
112 at  $0.82 \pm 1.28$  Ma, and hence interpreted as due to the Pleistocene-Holocene extensional phase. The other  
113 slickenfiber yielded a  $13.5 \pm 2.47$  Ma age, and interpreted as pre-thrusting, foreland flexure-related, Early-  
114 Middle Miocene extensional stage. However, this slickenfiber is crosscut by a second generation of  
115 slickenfibers with normal movement, indicating that the fault was later reactivated. The Quaternary seismic  
116 activity of the Val d'Agri fault systems is constrained by paleoseismological, historical, and instrumental data  
117 (e.g., Giano et al., 2000; Improta et al., 2017). The most destructive historical earthquakes occurred in 1857  
118 ( $M_w = 7.1$ ) with an uncertain source, which is either postulated to be the EAFS (Benedetti et al., 1988; Bello  
119 et al., 2022) or the MMFS (Maschio et al., 2005; Improta et al., 2010). Recent background seismicity is either  
120 of natural and/or man-made origin. Natural seismicity is consistent with the present NE-SW oriented  
121 extension (Maggi et al., 2009; Improta et al., 2017) and is primarily correlated with the southernmost  
122 segments of the MMFS (Valoroso et al., 2009; Improta et al., 2017). Induced events were triggered by  
123 oscillations in the water level of Pertusillo lake (Fig. 1c) or by wastewater re-injection during hydrocarbon  
124 extraction (Improta et al., 2017; Valoroso et al., 2023). The interaction between fluids and seismic activity in  
125 the study area is further evidenced by hydrogeochemical anomalies detected in a 400 m deep exploration  
126 well in the MMFS (i.e., Tramutola oil seep; Fig. 1d), which revealed significant short-term geochemical

127 variations before and after two moderate earthquakes in 1996 ( $M_L = 4.9$ ; Italiano et al., 2001) and 2004 ( $M_L$   
128 = 4.1; Colangelo et al., 2007). Since the 1860 earthquake, natural hydrocarbon emissions have been  
129 documented in the area surrounding this well (Colangelo et al., 2005).

130 Previous studies have examined syn-tectonic mineralizations exposed near Marsico Nuovo (Fig. 1c). They  
131 revealed the involvement of diagenetic fluids during the early orogenic shortening stage (Miocene; Iannace  
132 et al., 2012; Gabellone et al., 2013), of deep dolomitizing fluids likely originated from the Irpinina mélange  
133 during the late orogenic shortening stage (Pliocene; Iannace et al., 2012; Gabellone et al., 2013), and of warm  
134 fluids ( $130^{\circ}$ - $140^{\circ}$ ) during the post-orogenic extension (Pleistocene-Holocene; Mazzoli et al., 2004).



136 *Fig. 1. a) Structural-tectonic map of the Southern Apennines modified by Maggi et al. (2009) and Vitale & Ciarcia (2018),*  
137 *showing the location of panel-d. b) Tectono-stratigraphy succession with lithological description of tectonic units (Butler*  
138 *et al., 2004; Mazzoli et al., 2001; Patacca & Scandone, 2007; Palladino et al., 2023); c) Schematic cross-section (modified*  
139 *after Schirripa Spagnolo et al., 2024), whose trace is reported in panel-d; d) Structural-tectonic map of the Val d'Agri*  
140 *Basin (modified after Beaubien et al., 2023), showing sampling sites (blue-sky diamonds). EAFS: East Agri Fault System;*  
141 *MMFS: Monti della Maddalena Fault System.*

### 142 **3. Methods**

143 We collected 351 fault-related calcite mineralizations from fault zones pertaining either to the EAFS or to  
144 the MMFS, as well as 33 host rock samples representative of different tectonic units (Fig. 1c; Tables S1-  
145 S2). Far from the EAFS and MMFS fault zones, background veins with thickness suitable for micro-drilling  
146 are very rare, and generally oriented parallel to the bedding. Therefore, these veins are not statistically  
147 significant and were not sampled. Based on the amount of available material, we selected samples for the  
148 following analyses: 1) measurements of vein attitudes and of fault attitudes and kinematics; 2)  
149 microstructural observations of 42 thin sections by optical and cathodoluminescence microscopy; 3)  
150 carbon and oxygen isotope of 351 fault-related calcite mineralizations and 33 calcite host rocks using a  
151 ThermoFisher GasBench, requiring 100-150  $\mu\text{g}$  of sample; 4) carbonate clumped isotopes of 75 fault-  
152 related calcite mineralizations, using a Thermo Scientific Kiel IV-MAT253 system, which provides accurate  
153 results with only 400-800  $\mu\text{g}$  of sample (Muller et al., 2017). From clumped measurements ( $\Delta_{47}$ ; I-CDES)  
154 we derive precipitation temperatures ( $T_{\Delta_{47}}$ ; °C) and oxygen isotope compositions of parental fluids using  
155 calibrations of Anderson et al. (2021) and O'Neil et al. (1969), respectively ; 5) rare earth elements and  
156 yttrium (REY) of 9 fault-related calcite mineralizations and 3 host rock samples using a ICP mass  
157 spectrometer, requiring 5g of sample; 6)  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of 12 selected fault-related calcite mineralizations  
158 using a FINNIGAN MAT mass spectrometer requiring 20 mg of sample; 7) mineralogy of two clay-rich  
159 interbedded layers using a X-ray system equipped with silicon-strip detector, requiring 15g of sample.  
160 Detailed descriptions of the analytical methods are provided in the Supplementary Materials.

161



## 162 4. Results

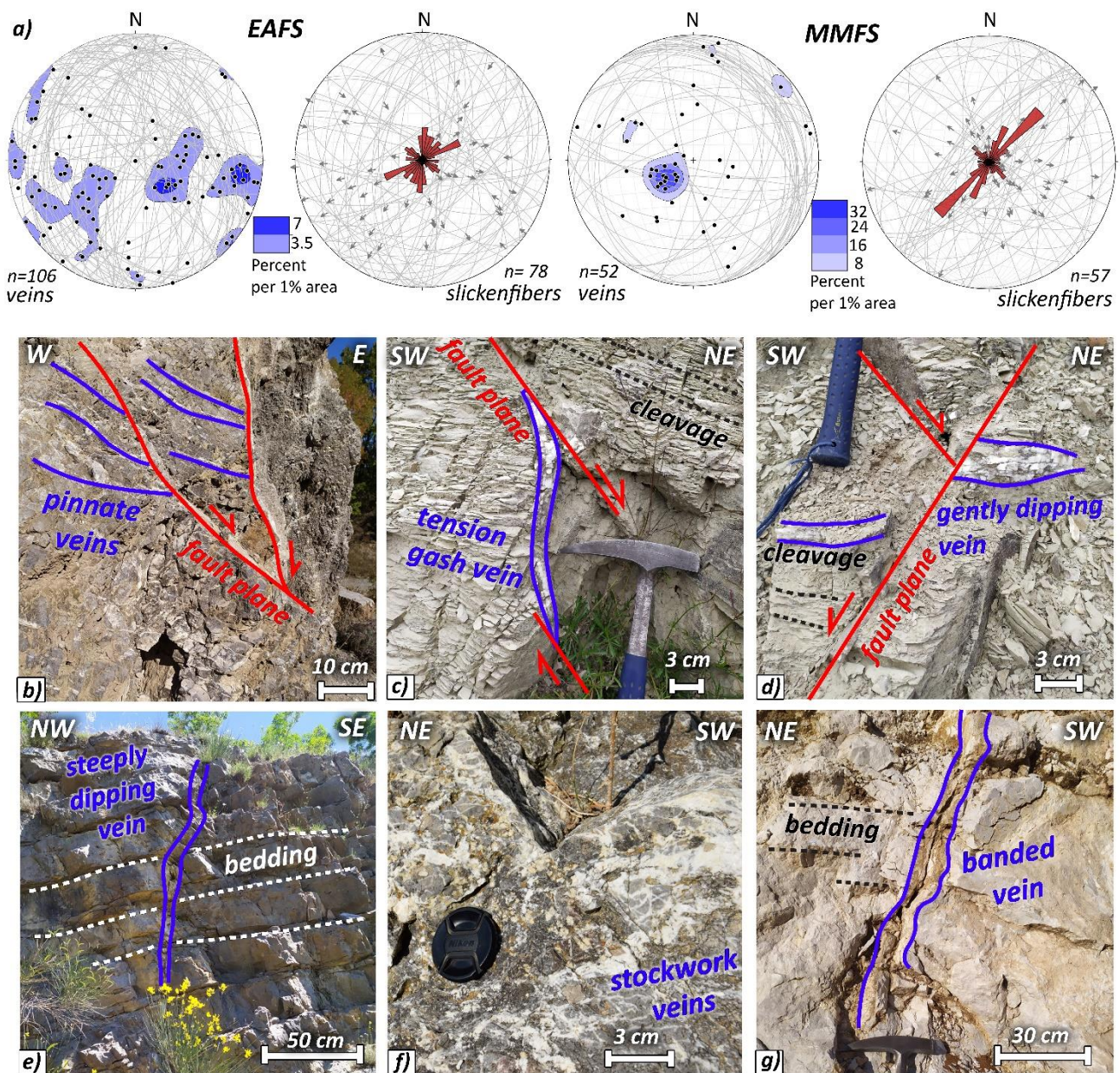
### 163 4.1. Meso-structural Observations

164 Among the 351 fault-related calcite mineralizations, we distinguish between those that are clear kinematic  
165 indicators and those that are not. In the Discussions section, we assign all mineralizations to the  
166 compressional or extensional tectonic phase with different degrees of confidence (Roberts & Holdsworth,  
167 2022). Mineralizations that are fault kinematic indicators are classified using the terminology of Van Der  
168 Pluijm & Marshak (2004) and include:

- 169 - slickenfibers on fault surfaces characterized by mineral growth lineations primarily indicating  
170 extensional shear sense, with a few exceptions showing compressional kinematics (Fig. 2a; Table S1).
- 171 - pinnate veins immediately adjacent to fault surfaces, forming an acute angle of 30°-45° relative to  
172 the fault plane and indicating extensional shear (Fig. 2b; Table S1).
- 173 - tension gashes characterized by sigmoidal shape along tension fractures, displaying contractional,  
174 strike-slip (both dextral and sinistral), or extensional shear (Fig. 2c; Table S1).

175 Other mineralizations, not associated with kinematic indicators, display a wide range of orientations (Fig. 2a),  
176 and are classified using geometric criteria. They include:

- 177 - gently dipping veins, with dip angles between 5° and 45° (Fig. 2a). Generally parallel to bedding or  
178 cleavage (Table S2), these veins are, in some cases, crosscut by extensional-transtensional faults (Fig.  
179 2d).
- 180 - steeply dipping veins, with dip angles between 50° and 90° (Fig. 2a, e). When bedding is recognizable,  
181 cut-off angles between vein and bed are > 50° (Table S2).
- 182 - stockwork veins, characterized by nonsystematic orientation, irregular shapes, and nonplanar array.  
183 These veins occur within pervasively fractured rocks, primarily at the footwall of the extensional-  
184 transtensional faults (Fig. 2f).
- 185 - banded veins, without any preferential attitude, and composed of 3 to 10 cm-wide white to brownish  
186 bands (Tables S1-2; Fig. 2g). These veins cut across bedding as well as extensional-transtensional  
187 faults and fractures (Fig. S1).



189

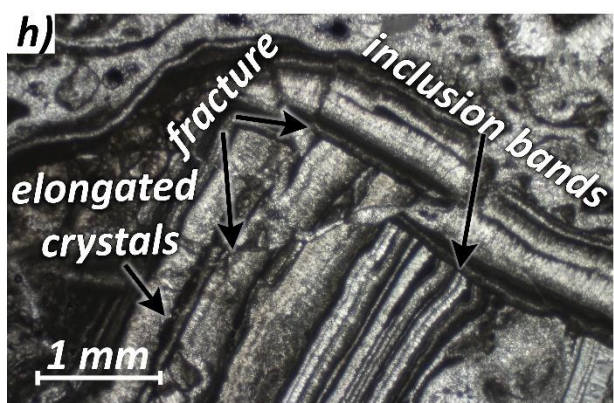
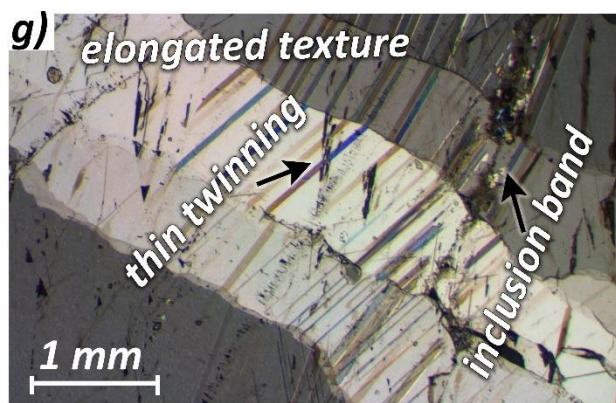
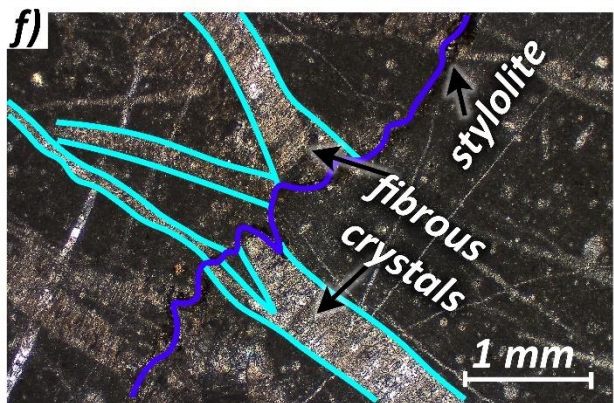
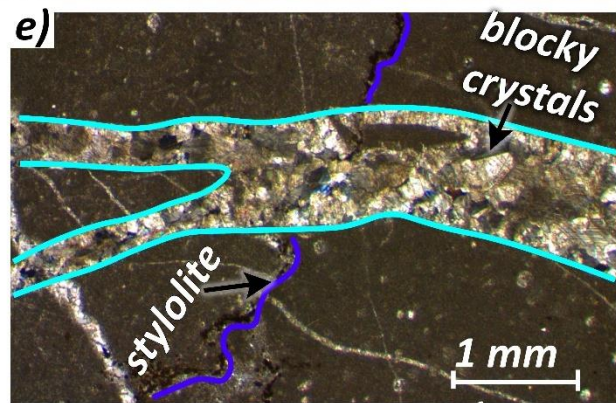
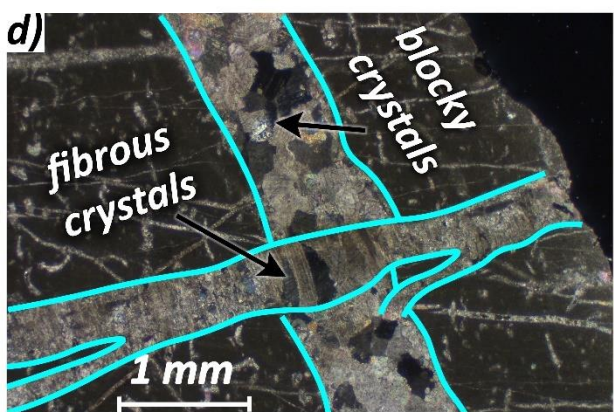
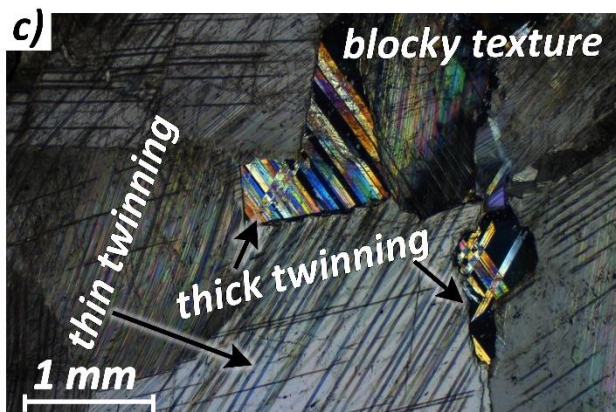
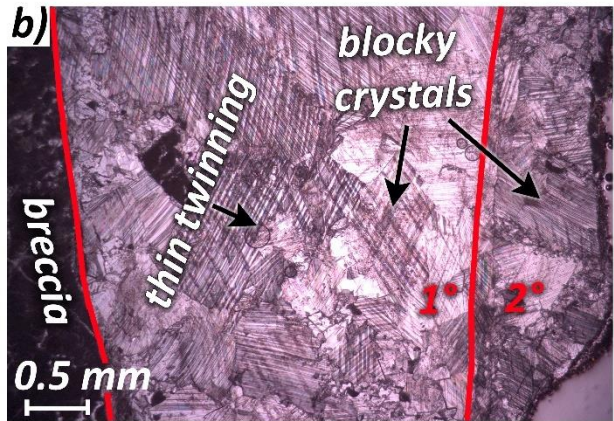
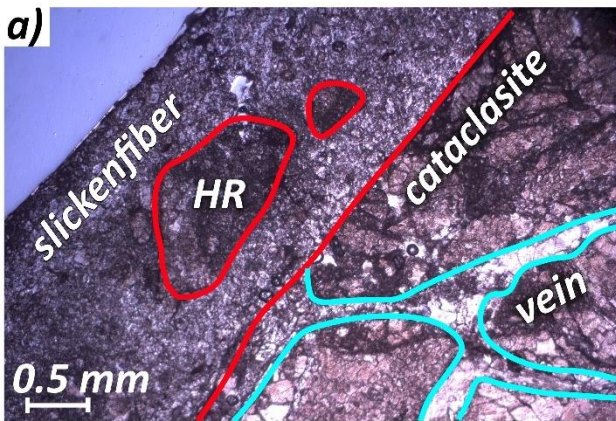
190 *Fig. 2. a) Schmidt projections (lower hemisphere) distinguished on EAFS and MMFS provenance, showing attitudes of*  
 191 *veins and fault planes with striae (grey), pole contour of veins (blue), and rose diagram of faults decorated by*  
 192 *slickenfibers (red). b) Outcrop picture (sampling site 10; Fig. 1) of pinnate veins, associated with extensional faults; c)*  
 193 *Outcrop picture (sampling site 6; Fig. 1) of sigmoidal tension gash vein, associated with extensional shear; d) Outcrop*  
 194 *picture (sampling site 7; Fig. 1) of gently dipping vein, almost parallel to cleavage and cut by an extensional fault; e)*  
 195 *Outcrop picture (sampling site 2; Fig. 1) of steeply dipping vein, with cut-off angles between vein and bed of ~70°; f)*  
 196 *Outcrop picture (sampling site 1; Fig. 1) of stockwork veins sampled at the footwall of a regional transtensional fault; g)*  
 197 *Outcrop picture (sampling site 5; Fig. 1) of banded vein.*

198 **4.2. Microstructures**

- 199 - Slickenfibers are characterized by blocky crystals and are frequently associated with coat breccias  
200 and cataclastic fault-rocks rich in veins (Fig. 3a). MAR29 and MAR8 samples exhibit crosscutting  
201 relationships, indicating two growth events (Fig. S2).
- 202 - Steeply to gently dipping and stockworks veins are composed of blocky calcite crystals with thin to  
203 thick twinning (Fig. 3). GIU22 and MA25 stockwork vein samples also show fibrous crystals (Figs. 3c,  
204 d). In some steeply and gently dipping veins (samples VA24, VA25, MA46, S10-39; Figs. 4 and S5),  
205 different crystal growth phases are identified. Furthermore, in samples including stylolites (Figs. 3d-  
206 f, S6), veins cut across them (sample VA17) or exhibit mutually crosscutting relationships (samples  
207 VA25, MA20, MA25, GIU22, and GIU33).
- 208 - Banded veins are characterized by blocky-elongated calcite crystals, with crystal growth competition,  
209 thin twinning, and inclusion bands parallel to vein walls (Fig. 3). Less commonly, alabastrine-like  
210 textures and open vugs are also present. Samples S2G1, S2G8, and S2G18 exhibit fractured crystals,  
211 indicating multiple growth phases (Fig. S1).

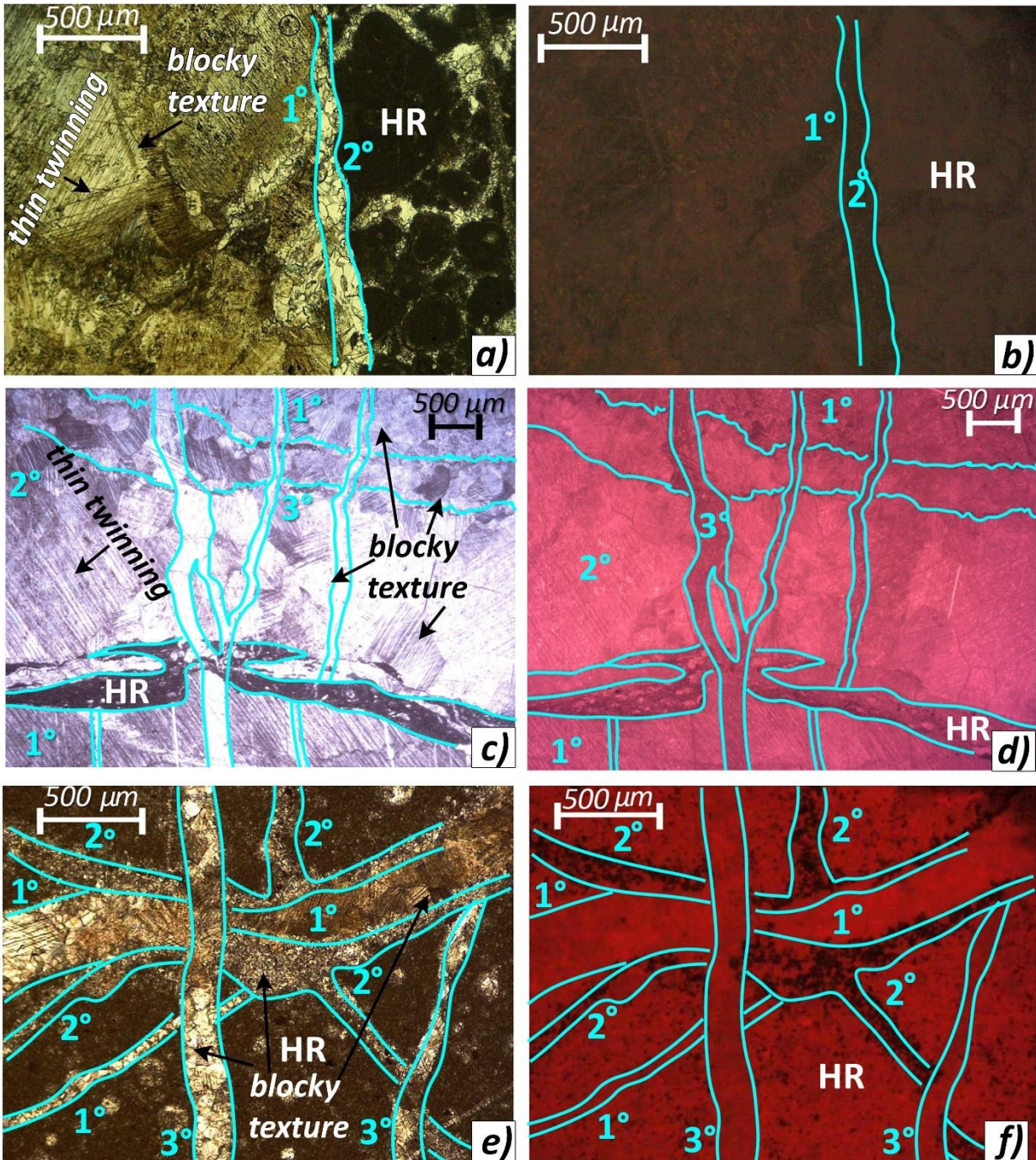
212 The cathodoluminescence colors of these syn-kinematic calcite mineralizations are lithology-dependent (Fig.  
213 4). Slickenfibers and veins in the Lagonegro Basin (Figs. 4e, f, S5) and Albidona Fm. (Figs. 4c, d, S4) show bright  
214 red luminescence, which is duller, homogeneous, or brighter than that of the host rock (except for sample  
215 MA52, which is completely not luminescent; Fig. S2). Conversely, syn-kinematic calcite mineralizations within  
216 the Apennine Carbonate Platform are non-luminescent (Figs. S1-2) or show dull red luminescence similar to  
217 that of the host rock (Figs. 4a, b, S3).

218



221 *Fig. 3* Optical images of: a) slickenfiber (MAR31 sample; Table S1), which coat cataclasite with veins; b)  
222 slickenfibers (MAR8 sample; Table S1) formed by two generations of blocky calcite crystals (1°, the oldest, and  
223 2°, the youngest); c) blocky texture with thin and thick twinning (S9-22 sample; Table S1); d) a vein with fibrous  
224 crystals which cut through a vein with blocky crystals (MA25 sample; Table S1); e) a vein with blocky crystals  
225 which cut through a stylolite (VA17 sample; Table S1); f) veins with fibrous crystals cut by stylolites (GIU22  
226 sample; Table S1); g) vein with elongated texture, thin twinning and inclusion bands (VA39 sample, Table S1);  
227 h) vein with fractured elongated crystals (S2G8 sample; Table S1). All images are taken with cross-polarized  
228 light except for panel a, for which plane-polarized light was used.

229



## OPTICAL MICROSCOPY

## CL MICROSCOPY

230

231 *Fig. 4 a), b) Optical and cathodoluminescence (CL) microscopy images of veins from the Apennine Carbonate Platform*  
 232 *(MA46 sample; Table S1), which show dull luminescence. c), d) Optical and CL microscopy images of veins from the*  
 233 *Albidona Fm., (S10-39 sample; Table S1), which show intense luminescence. e), f) Optical and CL microscopy images of*  
 234 *veins from the Lagonegro Basin unit (VA25 sample; Table S1), which show intense luminescence. HR: host rocks. Sky blue*  
 235 *numbers refer to different generations of calcite crystals, from the oldest (1) to the youngest (3). All optical images are*  
 236 *taken using plane-polarized light.*

### 237 4.3. Clay mineralogy

238 We carried out XRD analyses to further constrain the 1D burial and thermal models by Aldega et al. (2003)  
239 and Corrado et al. (2005) that indicate a maximum burial temperature of 140°-160°C for the Triassic-Jurassic  
240 limestones of the Lagonegro Basin due to a tectonic loading of 4-4.5 km. We analyzed two clay-rich layers  
241 interbedded within the Albidona Fm. (MAR42, Table S1; at site 6, Fig. 1c), and the Cretaceous limestones of  
242 the Apennine Carbonate Platform (MAR11, Table S1; at site 12, Fig. 1c) both overlying the Lagonegro Basin  
243 unit. The sample from the Albidona Fm. consists of kaolinite (61%), subordinate amounts of mixed layer illite-  
244 smectite (I-S, 20%), illite (18%), chlorite (1%), and traces of calcite and goethite (<1%). Mixed layers I-S mainly  
245 correspond to short-ordered structures (R1) with an illite content of 73% (Fig. S7), indicating the first stages  
246 of deep diagenesis with an approximate maximum temperature of 110°-120°C (e.g., Środoń, 1999). The XRD  
247 pattern also shows a small hump at  $5.2^{\circ}2\theta$  and aperiodic  $00l$  diffraction peaks interpreted as random ordered  
248 mixed layer I-S (R0).

249 The sample from the Cretaceous limestone of the Apennine Platform contains kaolinite (35%), illite (36%),  
250 two populations of mixed layers illite-smectite (26%), and low chlorite contents (3%). The most abundant  
251 population of I-S corresponds to short-ordered structures (R1) with an illite content of 78% (Fig. S7), whereas  
252 the second is composed of random ordered I-S (R0). This assemblage indicates a temperature range of 100°-  
253 110°C (Środoń, 1999).

254 Temperatures of 100°-120°C obtained for the Apennine Carbonate Platform and Albidona Fm. overlying the  
255 Lagonegro Basin unit are consistent with Aldega et al. (2003) and Corrado et al. (2005) who determined that  
256 the Miocene deposits at top of the Lagonegro succession (Fig. 1b) experienced maximum burial temperatures  
257 of 110°-120°C. Results from MAR11 and MAR 42 as well as from the Lagonegro succession (Aldega et al.,  
258 2003; Corrado et al., 2005) document that I-S formed during diagenetic overprint as a result of Miocene  
259 tectonic burial (Fig. S8).

260

#### 261 **4.4. Carbon, Oxygen, and Carbonate Clumped Isotopes**

262 The carbon and oxygen isotope compositions of veins and slickenfibers, along with those of the  
263 corresponding host rocks, are presented in Figure 5. The data from the Apennine Carbonate Platform are  
264 divided into two plots, separating samples from the EAFS and MMFS (Figs. 5c, d). The Apennine Carbonate  
265 Platform is extensively exposed along both the EAFS and the MMFS, whereas the Albidona Fm. and Lagonegro  
266 Basin rocks are predominantly exposed along the EAFS. Therefore, Figures 5a and 5b exclusively present data  
267 from the EAFS, with three exceptions indicated by circles in Figure 5a.

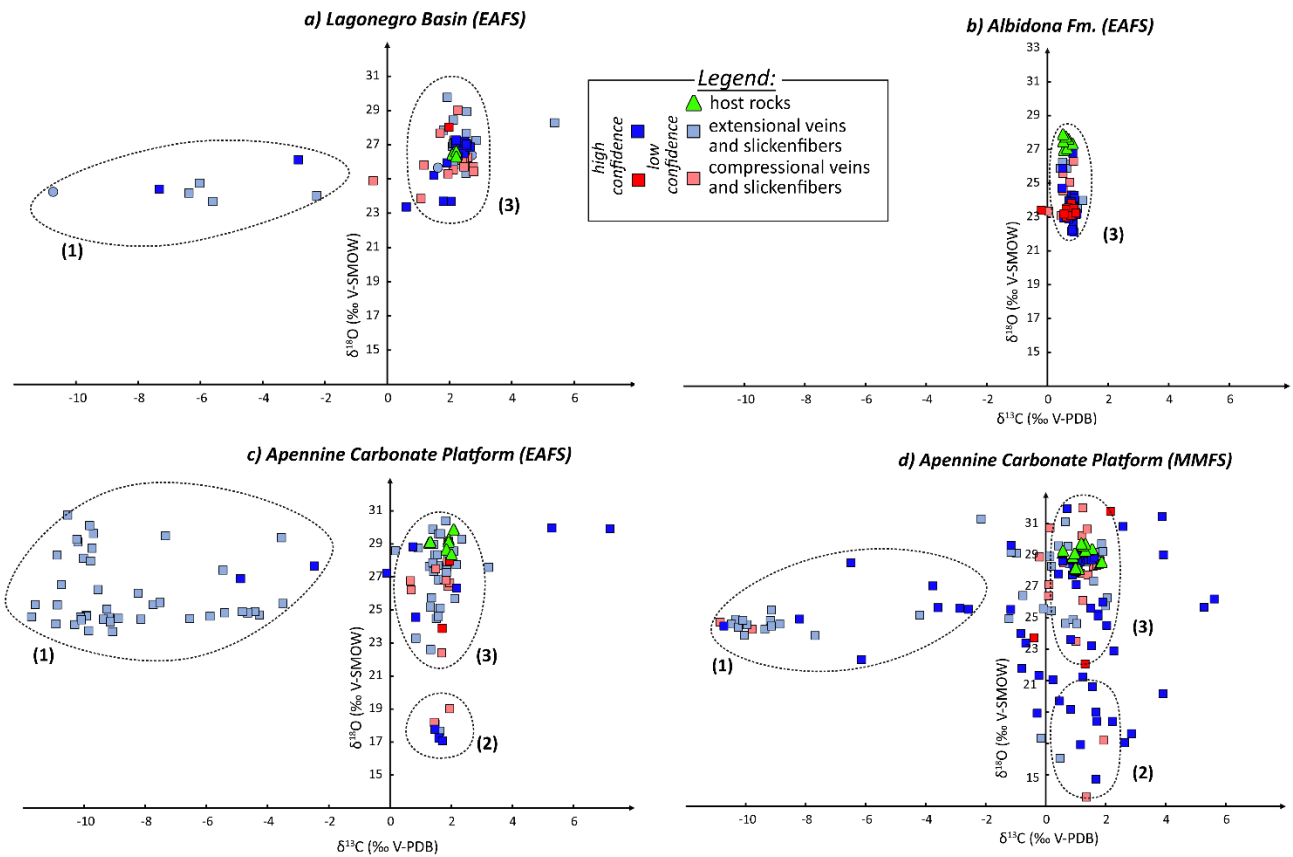
268 The Lagonegro Basin rocks are characterized by a  $\delta^{18}\text{O}$  values between +26.3 and +26.7‰ V-SMOW and by  
269  $\delta^{13}\text{C}$  values between +2.1 and + 2.3‰ V-PDB (Fig. 5a). The Albidona Fm. rocks show  $\delta^{18}\text{O}$  values between  
270 +26.9 and +27.9‰, and  $\delta^{13}\text{C}$  values between +0.5 and +0.9‰ (Fig. 5b). The Apennine Carbonate Platform  
271 rocks exhibit by  $\delta^{18}\text{O}$  values between +28.3 and +29.9‰, and  $\delta^{13}\text{C}$  values between +1.3 and +2.2‰ along the  
272 EAFS (Fig. 5c), and by  $\delta^{18}\text{O}$  values between +27.9 and +29.7‰, and  $\delta^{13}\text{C}$  values between +0.5 and +1.5‰  
273 along the MMFS (Fig. 5d). In Figure 5 we categorize veins and slickenfibers based on their  $\delta^{13}\text{C}$ . Samples with  
274 negative  $\delta^{13}\text{C}$  define cluster 1. Among the samples with positive  $\delta^{13}\text{C}$ , we define cluster 2 for veins and  
275 slickenfibers that have a  $\delta^{18}\text{O}$  more than 5‰ lower than the host rocks. Cluster 3 includes veins and  
276 slickenfibers with  $\delta^{18}\text{O}$  within 5‰ of the host rocks.

277 The Lagonegro Basin and the Albidona Fm. mineralizations show a relatively small variability and mostly  
278 belong to cluster 3 (Figs. 5a and b). The Apennine Carbonate Platform mineralizations show more scattered  
279  $\delta^{18}\text{O}$  values, falling into all three calcite clusters. Clusters 2 and 3 are well defined in samples from the EAFS  
280 (Fig. 5c), whereas they are not clearly separated in samples from the MMFS (Fig. 5d).

281

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283

284 *Fig. 5.  $\delta^{13}\text{C}$  vs.  $\delta^{18}\text{O}$  plots of calcite mineralizations collected from the Lagonegro Basin, the Albidona Fm., and the*

285 *Apennine Carbonate Platform (separated for EAFS and MMFS provenance). The circles in panel a indicate veins collected*

286 *from the Lagonegro Basin unit along the MMFS. In all panels, dashed circles indicate the different calcite clusters (1, 2,*

287 *and 3). High confidence data are for clear fault kinematic indicators, whereas low confidence data are for the other*

288 *mineralizations (see their tectonic interpretation in the Discussion section).*

289

290 We performed carbonate clumped isotopes analyses on 75 samples from all 3 clusters (Table S1). Figure 6

291 displays the calculated precipitation temperatures ( $T_{\Delta 47}$ ) vs. the calculated  $\delta^{18}\text{O}$  of the parental fluids. The

292 vertical grey lines represent the maximum temperatures experienced by the host rocks during burial. We

293 identify (Fig. 6a):

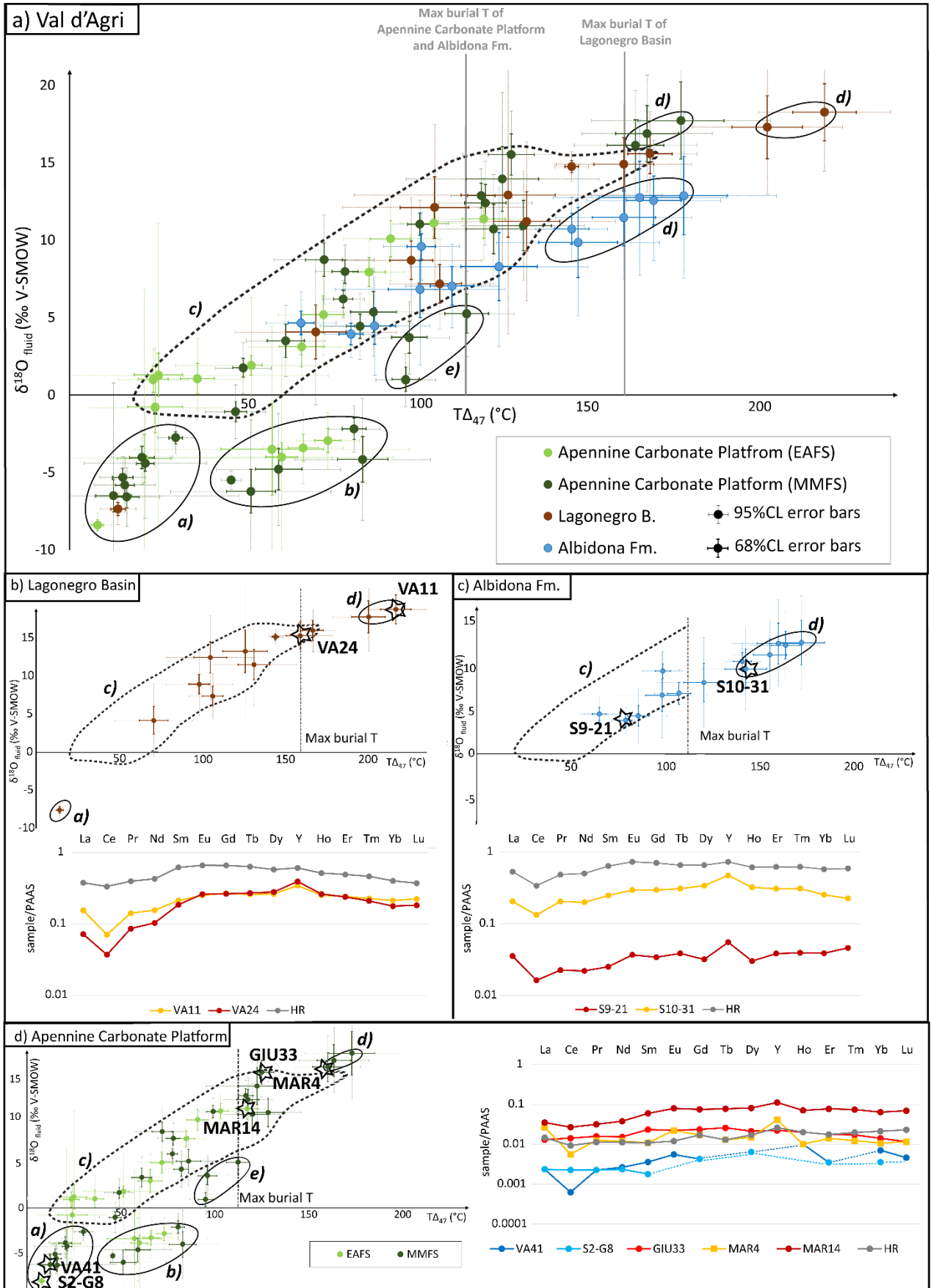
- 294 - Fluid group *a* characterized by parental fluids with  $\delta^{18}\text{O}$  between -2 and -10 ‰ and  $T_{\Delta 47}$  between
- 295 10° and 30°C. These data were obtained from calcite mineralizations of cluster 1;
- 296 - Fluid group *b* characterized by parental fluids with  $\delta^{18}\text{O}$  between -2 and -7 ‰ and  $T_{\Delta 47}$  between 40°
- 297 and 80°C, all obtained from calcite mineralization of cluster 2;

- 298 - Fluid group *c* characterized by parental fluids with  $\delta^{18}\text{O}$  between -1 and +16 ‰ and  $T\Delta_{47}$  between 30°  
299 and 160°C, lower than maximum burial temperature experienced by the host rocks (120°C for the  
300 Albidona Fm. and Apennine Carbonate Platform; 160°C for the Lagonegro Basin unit), obtained from  
301 calcite mineralizations of cluster 3;
- 302 - Fluid group *d* characterized by parental fluids with  $\delta^{18}\text{O}$  between +9 and +18 ‰ and  $T\Delta_{47}$  between  
303 150° and 220°C, higher than maximum burial experienced by the host rocks, obtained from calcite  
304 mineralizations of cluster 3;
- 305 - Fluid group *e* characterized by parental fluids with positive  $\delta^{18}\text{O}$  between 1 and 5 ‰,  $T\Delta_{47}$  between  
306 100° and 120°C, obtained from calcite mineralizations of cluster 2.

307

308

309



311 *Fig. 6. a)  $\delta^{18}\text{O}_{\text{fluid}}$  vs.  $T_{\Delta 47}$  plot for the fault-related calcite mineralizations, showing the recognized fluid groups (a, b, c,*  
312 *d, and e). The vertical grey lines represent the maximum temperatures experienced by the host rocks during burial. These*  
313 *burial temperatures are obtained by clay mineralogy as presented in Section 4.3. Error bars of  $\delta^{18}\text{O}_{\text{fluid}}$  and  $T_{\Delta 47}$  are*  
314 *reported at both 68% (solid lines) and 95% (dotted lines) confidence level (CL; Table S1). b), c), d)  $\delta^{18}\text{O}_{\text{fluid}}$  (‰ V-SMOW)*  
315 *vs.  $T_{\Delta 47}$  (°C) plot and Post Archean Australian shale (PAAS)-normalized REY trend (rare Earth elements + yttrium) of*  
316 *samples from Lagonegro Basin (b), Albidona Fm. (c), and Apennine Carbonate Platform (d). Stars and letters indicate the*  
317 *samples with the associated fluid group for which we analyzed REY elements. HR: host rock.*

#### 318 **4.5. Rare-earth elements + Yttrium and Strontium isotopes**

319 To provide additional constraints on parental fluids through REY and Sr isotope analyses, we selected samples  
320 representative of fluid groups a (VA41 and S2G8), c (MAR14, VA24, GIU33, and S9-21) and d (S11-11, MAR4,  
321 VA11, S10-15, S10-38, and S10-31; Table1).

322 Figure 6 compares the REY pattern of mineralizations normalized to PAAS (Post Archean Australian Shales,  
323 McLennan, 1989), with those of the corresponding host rocks. REY patterns of all veins and slickenfibers from  
324 the Lagonegro Basin unit and the Albidona Fm. mimic those of the host rocks, which are characterized by  
325 negative Ce and positive Y anomalies (Figs. 6b and c). Mineralizations are depleted in REE and are  
326 characterized by more pronounced Ce and Y anomalies than the host rocks (Table 1). This pattern is more  
327 evident for samples from fluid group c than those from group d (Table 1). The three analyzed fluid groups of  
328 the Apennine Carbonate Platform show different patterns of REY (Fig. 6d): group a is very depleted in REY  
329 elements; group c mimics the flat REY trend of the host rock and group d shows evident negative Ce as well  
330 as positive Eu and Y anomalies (Table 1).

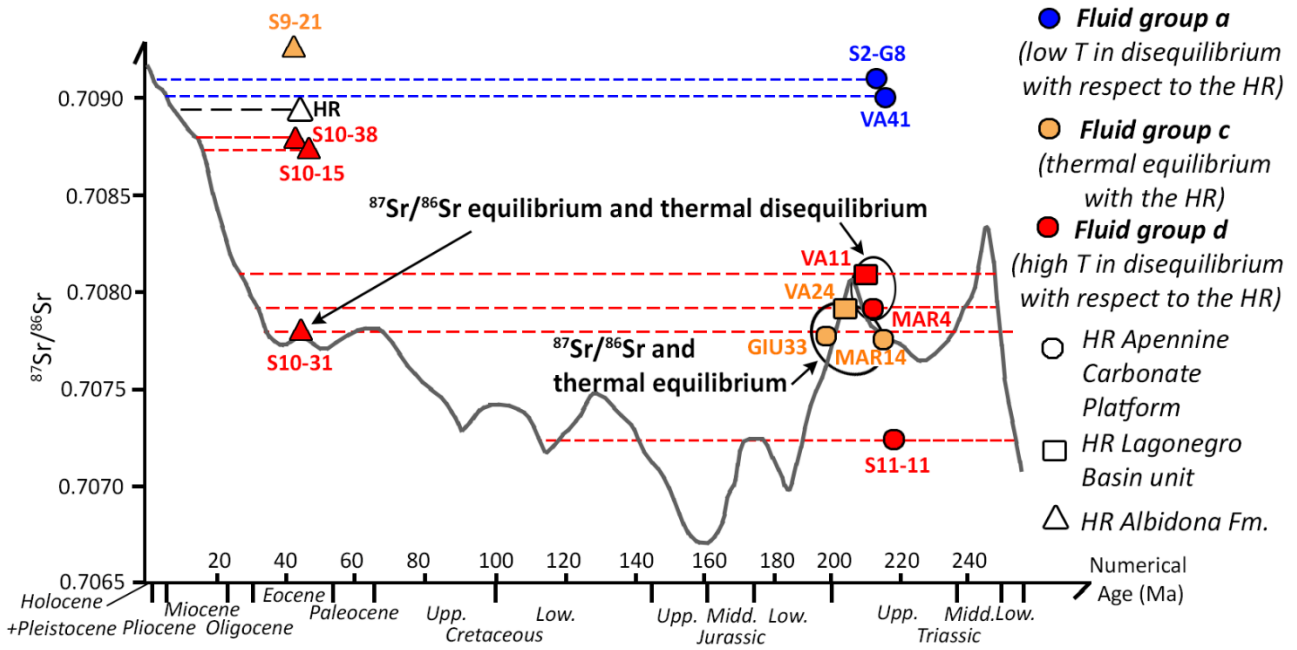
331 Since the calcite veins and slickenfibers have low Rb concentrations (Table S3), the contribution of radiogenic  
332  $^{87}\text{Sr}$  is virtually negligible. Therefore, the measured  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio primarily reflects fluid interactions with  
333 rocks of specific stratigraphic ages and isotopic signatures. However, to account for the possible  
334 contamination of the Sr signature since the interaction with radiogenic fluids derived from dehydrated clays,  
335 we also calculated the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of the clay-rich Albidona Fm. rocks. This analysis revealed a particularly  
336 radiogenic ratio of 0.70895 for these rocks (Fig. 7).

337 In Figure 7, we compare the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios of mineralizations with the expected values of the age of the  
338 corresponding host rocks (McArthur et al., 2001). Except for sample S9-21, which is characterized by a  
339 particularly high  $^{87}\text{Sr}/^{86}\text{Sr}$  value, samples from fluid group *c* have  $^{87}\text{Sr}/^{86}\text{Sr}$  consistent with the stratigraphic  
340 age of the host rocks. Samples of fluid groups *a* and *d* display Sr isotope disequilibrium with respect to the  
341 host rocks. Specifically, fluid group *a* exhibits a Pliocene-Pleistocene isotopic signature, in disequilibrium with  
342 the Late Triassic age of the host rocks. Fluid group *d* shows a Miocene isotopic signature in disequilibrium  
343 with the Eocene age of the host rocks (samples S10-15, S10-38), and an Early Cretaceous/Middle-Early  
344 Jurassic isotopic signature in disequilibrium with the Late Triassic age of the host rocks (S11-11 sample).  
345 Conversely, samples VA11, S10-31 and MAR4 of fluid group *d* show  $^{87}\text{Sr}/^{86}\text{Sr}$  consistent with the host rocks  
346 (Fig. 7).

347

Table 1. Geochemical data for representative samples of different fluid groups

sample	Site & fault system	unit	mineralization type	$T_{\Delta 47}$ (°C) ( $\pm 95\%CL$ )	$\delta^{18}O_{fluid}$ (‰ V-SMOW) ( $\pm 95\%CL$ )	$^{87}Sr/^{86}Sr$ ( $\pm se$ )	Sr (ppm)	$\Sigma REY$ (ppm)	La*/Lu*	Y/Ho	Eu/Eu*	Ce/Ce*	Y/Y*	
fluid group a	VA41	11 MMFS	L. Triassic dolostones of Ap. Carb. Plat.	banded vein	13 $\pm$ 36	-6.50 $\pm$ 7.26	0.708994 $\pm$ 0.000009	23	0.35	0.51				
	S2G8	5 EAFS	L. Triassic dolostones of Ap. Carb. Plat.	banded vein	8 $\pm$ 4	-8.40 $\pm$ 0.82	0.709101 $\pm$ 0.000008	63	0.44					
fluid group c	MAR 14	12 EAFS	L. Triassic-E. Cretaceous limestones of Ap. Carb. Plat.	Gently dipping (bed parallel) vein	120 $\pm$ 13	11.36 $\pm$ 2.63	0.707756 $\pm$ 0.000010	167	9.41	0.51	42.8	1.18	0.74	1.46
	GIU 33	18 MMFS	M. Jurassic limestones of Ap. Carb. Plat.	gently dipping veins	128 $\pm$ 14	15.53 $\pm$ 2.82	0.707768 $\pm$ 0.000012	258	3.27	1.13	30	0.95	1.03	1.07
	VA24	2 EAFS	L. Triassic limestones of Lagonegro B.	extensional slickenfibres	168 $\pm$ 14	15.58 $\pm$ 2.72	0.707899 $\pm$ 0.000007	1628	25.20	0.39	40.7	1.16	0.45	1.45
	S9- 21	6 EAFS	Eocene Albidona Fm.	compressional slickenfibers	82 $\pm$ 8	3.93 $\pm$ 1.51	0.709246 $\pm$ 0.000008	1431	5.78	0.77	50	1.24	0.53	1.78
fluid group d	S11- 11	10 MMFS	L. Triassic dolostones of Ap. Carb. Plat.	transtensional slickenfibers	177 $\pm$ 26	17.73 $\pm$ 5.20	0.70723 $\pm$ 0.000009							
	S10- 31	7 EAFS	Eocene Albidona Fm.	extensional slickenfibers	147 $\pm$ 24	9.84 $\pm$ 4.73	0.707785 $\pm$ 0.000012	1296	45.53	0.91	40	1.09	0.65	1.43
	MAR 4	23 MMFS	L. Triassic-E. Cretaceous limestones of Ap. Carb. Plat.	pinnate vein	164 $\pm$ 18	16.13 $\pm$ 3.54	0.707907 $\pm$ 0.000009	211	3.36	2.31	110	1.59	0.26	3.25
	VA11	1 EAFS	Late Triassic limestones of Lagonegro B.	transtensional slickenfibers	219 $\pm$ 19	18.27 $\pm$ 3.82	0.708066 $\pm$ 0.000009	513	32.07	0.69	37.2	1.04	0.46	1.34
	S10- 15	7 EAFS	Eocene Albidona Fm.	extensional slickenfibres	169 $\pm$ 19	12.56 $\pm$ 3.88	0.708745 $\pm$ 0.000008							
	S10- 38	7 EAFS	Eocene Albidona Fm.	transtensional slickenfibers	178 $\pm$ 27	12.87 $\pm$ 5.36	0.708776 $\pm$ 0.000012							



350

351 *Fig. 7.  $^{87}\text{Sr}/^{86}\text{Sr}$  values of fault-related calcite mineralizations compared with the Sr isotope evolution of seawater by*  
 352 *McArthur et al. (2001). The age of the host rocks is from Palladino et al. (2023). Samples are distinguished by the host*  
 353 *rock (HR) lithology and identified fluid groups (Table 1). The  $^{87}\text{Sr}/^{86}\text{Sr}$  value of the Albidona Fm. rock are shown in the*  
 354 *plot to highlight the influence of radiogenic strontium on syn-tectonic calcite mineralizations.*

355

## 356 5. Discussion

### 357 5.1. Parental fluid origin and evolution

358 Calcite mineralizations of cluster 1 is characterized by  $\delta^{13}\text{C}$  values comprised between -2 ‰ and -12  
359 ‰ (Fig. 5), indicating a contribution of light carbon derived by oxidation of organic matter (e.g., Sharp, 2017).  
360 This cluster precipitated from fluids with a  $\delta^{18}\text{O}$  between -2 and -10‰, and a  $T_{\Delta_{47}}$  between 10° and 30°C  
361 (group *a*; Fig. 6a). These values are consistent with the present-day composition, of meteoric water from the  
362 Val d'Agri area (i.e.,  $\delta^{18}\text{O}$  -8‰, and T between 18° and 30°C; Italiano et al., 2001). Thus, we interpret that  
363 calcite of cluster 1 precipitated from meteoric waters interacting with shallow soil organic carbon before  
364 their infiltration at depth.

365 Calcite mineralizations of cluster 2 show  $\delta^{13}\text{C}$  values similar to that of their host rocks, and  $\delta^{18}\text{O}$  values  
366 mostly between +14 ‰ and +21 ‰ (Fig. 5), in disequilibrium with respect to the host rocks. Calculated  
367 parental fluid compositions of this cluster have  $\delta^{18}\text{O}$  values between -2 and -7 ‰, hence retaining a meteoric  
368 water signature, but with a  $T_{\Delta_{47}}$  between 40° and 80°C (group *b*; Fig. 6a). Since the dissolved inorganic carbon  
369 is a trace component of water, the carbon isotope ratios can be rapidly buffered by the host rocks, even with  
370 a short residence time. Conversely, oxygen, a major component of water, requires a longer residence time  
371 or low water/rock ratios to be buffered (Sharp, 2017). Therefore, we conclude that the fluid source of calcite  
372 cluster 2 consists of meteoric water that infiltrated at depth and warmed up with a limited isotope exchange  
373 with the host rocks. We observe that samples S11-5, MAR27, and MAR30 (Table S1), from the calcite cluster  
374 2, precipitated from fluids with  $\delta^{18}\text{O}$  values between 1 and 5 ‰, and  $T_{\Delta_{47}}$  between 100° and 120°C (group *e*;  
375 Fig. 6a). Since these high temperatures are apparently at odds with the relatively low  $\delta^{18}\text{O}$  values, we  
376 interpret them as due to the mixing of warm fluids ( $T \gg 100^\circ\text{C}$ ) with colder, infiltrated meteoric water.

377 Calcite cluster 3 exhibits  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  values almost equal to the host rocks (Fig. 5). The calculated  
378  $\delta^{18}\text{O}$  values of parental fluids  $> -1$  ‰ (groups *c* and *d* in Fig. 6a) indicate either a low fluid/rock ratio, or a long-  
379 term fluid-host rock interaction with higher fluid/rock ratios (Sharp, 2017). For these fluids, we observe a  
380 wide range of  $T_{\Delta_{47}}$  and oxygen isotope values, demonstrating a wide variability in water-rock interaction,  
381 and/or in the mixing of different end-member fluids. The precipitation temperatures of 200°-220°C of some  
382 mineralizations from the Lagonegro Basin unit as well as 150°-180°C of some mineralizations from the



383 Apennine Carbonate Platform/Albidona Fm. (Fig. 6a) are significantly higher than the maximum burial  
384 temperature experienced by the host rocks (140°-160°C for the Lagonegro Basin unit and 110°-120°C for  
385 Albidona Fm. and Apennine Carbonate Platform, estimated by the thermal modeling of Aldega et al., 2003  
386 and Corrado et al., 2005 and integrated with our new XRD data presented above).

387 Summarizing, parental fluids of calcite cluster 3 are either in thermal equilibrium (group *c*) or  
388 disequilibrium (group *d*) with respect to the maximum burial temperatures of the host rocks. The sources of  
389 fluid group *c* could include: meteoric waters isotopically buffered and in thermal equilibrium with the host  
390 rocks; water expelled during the smectite to illite transformation reaction; and formation waters trapped  
391 within shaly interlayers and expelled due to compaction (e.g., Smeraglia et al., 2020). In contrast, the high  
392 temperatures of fluid group *d* require a source deeper than the host rocks. By assuming a geothermal  
393 gradient of 30°C/km, and a surface temperature of 10°C, as proposed by Aldega et al. (2003) and Corrado et  
394 al. (2005), the  $T_{\Delta_{47}}$  of 220°C from mineralizations within the Lagonegro unit is consistent with an origin of  
395 parental fluids from depths of ~ 7 km. This corresponds to the burial depths of the Apulian Carbonate Platform  
396 both during orogenic and post-orogenic phases. Indeed, considering a post-orogenic exhumation of ~4-5 km  
397 (Aldega et al., 2003) for the top of the Apulian Carbonate Platform, which is currently located at depths of 2-  
398 3 km beneath the Val d'Agri Basin (Butler et al., 2004; Hager et al., 2021), its maximum can be estimated to  
399 be approximately 6-8 km. Therefore, fluids with temperatures ranging from 150°-180°C that permeated the  
400 Albidona Fm. and Apennine Carbonate Platform probably derived from rocks at a depth of ~4.5-5.5 km. They  
401 could be derived from the Lagonegro Basin unit or the Apulia Carbonate Platform, depending on the extent  
402 of rock exhumation at the time of calcite precipitation.

403 These interpretations are also supported by REY and Sr isotope. Samples of calcite cluster 1 are strongly  
404 depleted in REY (Fig. 6d), consistent with precipitation from meteoric water (fluid group *a*; Mcleannan, 1985).  
405 On the other hand, most samples of calcite cluster 3 are depleted in LREE elements ( $La^*/Lu^* < 0.8$ ), are  
406 characterized by moderately negative Ce and positive Y anomalies, and show a Y/Ho ratio between 30 and  
407 50 (Table 1). This is consistent with seawater inheritance (Mcleannan, 1985; Tostevin et al., 2016) from the  
408 sedimentation environment of the carbonate host rocks and with a long-term fluid-rock interaction. The lack  
409 of Ce and Y anomalies in samples GIU33, S10-31, and VA11 of calcite cluster 3 may be due to contamination

410 with clay minerals. Clays containing high concentrations of REY with an almost flat trend (Fig. 6) would dilute  
411 any possible anomalies (e.g., Tostevin et al., 2016). Additionally, the MAR4 sample (fluid group *d*) is enriched  
412 in LREE ( $La^*/Lu^*=2.31$ ) and shows a pronounced Eu positive anomaly and  $Y/Ho = 110$  (Table 1), likely resulting  
413 from thermochemical sulphate reduction experienced by hydrocarbon-rich fluids (Jiang et al., 2015). This is  
414 consistent with the proposed origin for fluid group *d* from the Apulian Carbonate Platform, where the  
415 hydrocarbons are trapped.

416 Furthermore, samples S10-31, VA11, and MAR4 from fluid group *d* and samples VA24, GIU33, and MAR14  
417 from fluid group *c* show a Sr isotope signatures consistent with their host rocks, whereas samples S10-38,  
418 S10-15, S11-11 from fluid group *d* and samples S2-G8, VA41 from fluid group *a* are not consistent with the  
419 stratigraphic age of the host rocks (Fig. 7). Sample S9-21 (group *c*) shows anomalously high  $^{87}Sr/^{86}Sr$  values,  
420 which we interpret as due to a long-term interaction of parental fluid with the clay-rich Albidona Fm. host  
421 rock rich in radiogenic Sr (Fig. 7). The Pliocene-Pleistocene Sr isotope signatures of samples from fluid group  
422 *a* (Fig. 7) are consistent with precipitation from meteoric fluids, which chemically interacted with shallow  
423 Pliocene-Pleistocene continental deposits (0.70907-0.70910; Mancini et al., 2007). The Sr isotope ratios of  
424 samples from fluid group *c*, consistent with the stratigraphic age of the host rocks, confirm the host-rock  
425 buffering of these fluids (Fig. 7). The samples of group *d*, precipitated from hot fluids in thermal disequilibrium  
426 with the host rocks, suggest a fluid origin deeper than the host rocks. Consequently, we interpret the  
427 Miocene-age Sr isotope signature of samples S10-15 and S10-38 (Fig. 7) as due to the fluid interaction with  
428 the Irpinia mélange or the top of the Apulian Carbonate Platform, and the Early Cretaceous/Middle-Early  
429 Jurassic Sr isotope signature of sample S11-11 (Fig. 7), as due to the interaction with the middle portion of  
430 the Apulian Carbonate Platform (Fig. 1b). Possibly, the apparent Sr isotope ratio of samples VA11, MAR4, and  
431 S10-31 (fluid group *d*) consistent with the stratigraphic age of the host rocks (Fig. 7) could be due to the  
432 interaction with Late Triassic dolostones of the Apulian Carbonate Platform.

433

## 434 **5.2. Tectonic interpretation of fault-related mineralizations**

435 The studied mineralizations, sampled along the high-angle post-orogenic extensional faults pertaining either  
436 to the EAFS or to the MMFS, can be attributed with different levels of confidence to orogenic compressional  
437 or post-orogenic extensional tectonics (Table 2). Calcite samples forming kinematic indicators as  
438 slickenfibers, sigmoidal tension gashes, and pinnate veins are interpreted with high confidence as syn-  
439 kinematic (Roberts & Holdsworth, 2022). Based on their kinematics, these samples can be confidently  
440 associated with orogenic compressional or post-orogenic extensional faulting. On the contrary, the tectonic  
441 origin of steeply and gently dipping veins as well as banded and stockwork veins, which are not associated  
442 with straightforward kinematic indicators, can interpret with a low confidence level.

443 The attitude of steeply dipping veins is consistent with the shear sense of nearby post-orogenic extensional  
444 faults, suggesting a kinematic relationship, and is compatible with the present-day extensional stress regime  
445 (Mariucci and Montone, 2020). Furthermore, these veins are not cut or displaced by structures associated  
446 with orogenic compressional phases, reinforcing their association with the post-orogenic extensional phase.  
447 Hence, we interpret them as products of the post-orogenic extensional phase. The banded veins, which cut  
448 across bed interfaces and extensional faults/fractures (Fig. S1), likely formed during the youngest crack and  
449 seal episodes of the post-orogenic extension (Uysal et al., 2011). The gently dipping veins have a geometry  
450 incompatible with the kinematics of high-angle extensional faults and are generally subparallel to bedding.  
451 In some cases, these veins are also cut by high-angle extensional faults (Fig. 2d). Since this type of vein is  
452 frequent in fold-and-thrust belts and is generally associated with the early compressive phases and related  
453 fluid overpressures (Jessell et al., 1994; Sibson, 2020), we interpret them as products of the orogenic  
454 compressional phase. Stockwork veins are generally related to rapid fluid depressurization of pressurized  
455 fluids during faulting (Sibson, 2000; Bons et al., 2022). Here, we interpret them as due to compressional  
456 tectonics, when observed at the footwall of reverse faults, or to extensional tectonics when observed at the  
457 footwall of extensional-transtensional faults. Supporting this interpretation, at the microscale, stockwork  
458 veins observed near thrust faults (samples GIU22, GIU33, and VA25; Tables S1-S2) commonly display mutual  
459 crosscutting relationships with seismogram-like stylolites, interpreted as syn-orogenic by Manniello et al.

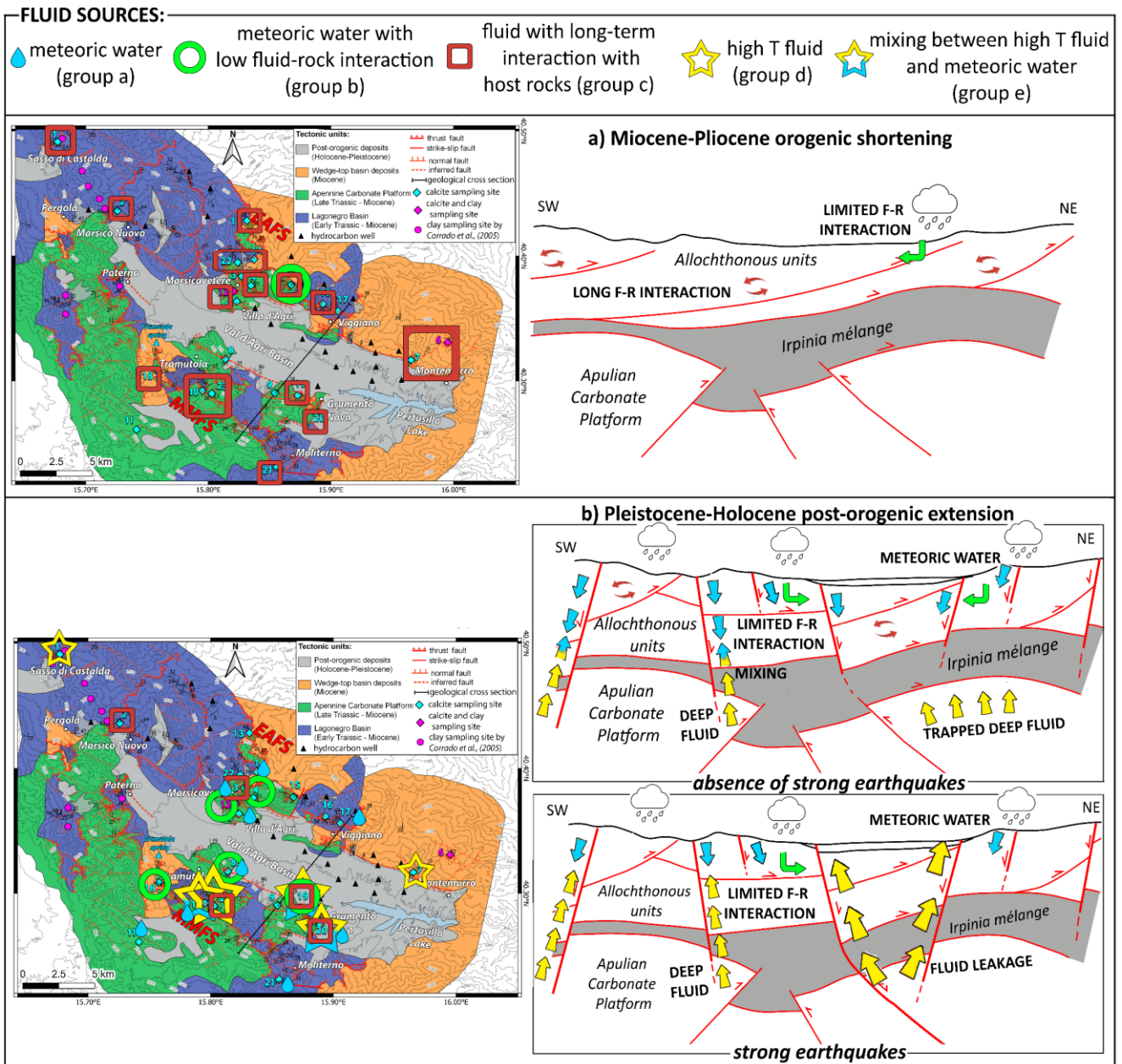
460 (2023), whereas those sampled near extensional faults (sample VA17; Table S1) cut across such stylolites (Fig.  
461 3).

462 In the Val d'Agri, both post-orogenic extensional and orogenic compressional faulting was preceded by rift-  
463 related and foreland flexure-related extensional faulting (Mazzoli et al., 2001; Patacca & Scandone, 2007;  
464 Schirripa Spagnolo et al., 2024). Thus, in principle, we cannot entirely rule out the possibility that some of  
465 the steeply dipping veins are inherited from such pre-orogenic tectonic stages. However, we consider them  
466 to be too rare to affect the robustness of our fluid circulation model, for the following reasons: 1) background  
467 veins, far from post-orogenic extensional faults, are very rare and generally parallel to bedding (i.e., not  
468 associated with extensional faulting); 2) the rare Mesozoic syn-sedimentary faults, recognized only in the  
469 Lagonegro Basin unit on stratigraphic grounds, were only weakly reactivated under contraction (Mazzoli et  
470 al., 2001); 3) no double inversion (first contractional and then extensional) of pre-orogenic faults has been  
471 recognized or described in the literature.

472

Table 2. Synthesis of structural-geochemical results characteristic of each fluid group

Fluid group and source	Tectonic unit and fault system	Myneralization type associated with post-orogenic tectonics	Myneralization type associated with orogenic tectonics	Micro-structural assemblage	Calcite cluster	$\delta^{18}\text{O}_{\text{fluid}}$ (‰ V-SMOW)	$T_{\Delta 47}$ (°C)	REY trend and anomalies	$^{87}\text{Sr}/^{86}\text{Sr}$ ratio
(a) Meteoric water	Lagonegro Basin and Apennine Carbonate Platform (both along EAFS and MMFS)	-banded veins (low confidence) - extensional slickenfibers (high confidence) -steeply dipping veins (low confidence)	-	Blocky and elongated texture, thin twinning, no luminescence	(1)	-3 to -9	10-30	Depleted in REY elements	Pliocene-Pleistocene Sr isotope signature
(b) Meteoric water with low residence time	Apennine Carbonate Platform (both along EAFS and MMFS)	-steeply dipping veins (low confidence) - extensional slickenfibers (high confidence)	- gently dipping (bed parallel) veins (low confidence)	Blocky texture, thin twinning, no luminescence	(2)	-3 to -7	50-80	-	-
(c) Host rock-buffered fluid (long residence time)	Albidona Fm., Lagonegro Basin, and Apennine Carbonate Platform (both along EAFS and MMFS)	- extensional slickenfibers (high confidence) - extensional tension gash (high confidence) - steeply dipping veins (low confidence)	- gently dipping (bed parallel) veins (low confidence) -compressional slickenfibers (high confidence) - compressional tension gash (high confidence) -stockwork veins near compressional faults (low confidence)	Blocky texture, thin and thick twinning; no luminescence or luminescence with same intensity of the HR one	(3)	-1 to 16	70-120 for Apennine Carbonate Platform and Albidona Fm., 100-170 for Lagonegro Basin unit.	moderate Ce, Y anomalies; Y/Ho <50	Sr isotope signature in equilibrium with respect to the stratigraphic age of the host rocks
(d) High T fluid (Apulian Carbonate Platform origin)	Albidona Fm., Lagonegro Basin, and MMFS Apennine Carbonate Platform	- pinnate veins (high confidence) - stockwork veins near extensional faults (low confidence) - extensional slickenfibers (high confidence)		Blocky texture thin and thick twinning, no luminescence or luminescence more intense than HR	(3)	10 to 18	200-220 for Lagonegro Basin unit  150-180 for Apennine Carbonate Platform and Albidona Fm.	moderate Ce, Y anomalies; Y/Ho > 100	Late Triassic, Early Cretaceous/ Middle Jurassic, Miocene Sr isotope signature
(e) Mixing between high T fluids and meteoric water	MMFS Apennine Carbonate Platform	- extensional slickenfibers (high confidence)		Blocky texture; thin twinning; no luminescence	(2)	1 to 5	100-120	-	-



474

475 *Fig. 8. Structural map of the Val d'Agri Basin (to the left) with spatial distribution of parental fluid groups involved during*  
 476 *the Miocene-Pliocene orogenic (panel a) or the Pleistocene-Holocene post-orogenic (panel b) tectonic phases. Schematic*  
 477 *models of fluid flow/distribution are provided for each phase. Notice that during the post-orogenic phase, fluid*  
 478 *circulation differs during strong earthquakes.*

### 479 5.3. Spatiotemporal fluid circulation

480 Based on the results discussed above, and synthesized in Table 2, mineralizations associated with orogenic  
481 shortening precipitated from fluids in chemical and thermal equilibrium with the host rocks (group *c*; Table  
482 2). This indicates a compartmentalized fluid-circulation system (e.g., Beaudoin et al., 2023) involving  
483 intraformational fluids and/or meteoric fluids completely buffered by the host rocks during the orogenic  
484 compressional tectonics (Fig. 8a). This scenario is consistent with the interpretation of Gabellone et al. (2013)  
485 for calcite veins from the Val d'Agri Basin, and with other case studies from the central Italian Apennines  
486 (e.g., Smeraglia et al., 2020; Curzi et al., 2024) and other fold-and-thrust worldwide belts (e.g., Beaudoin et  
487 al., 2023).

488 Mineralizations associated with post-orogenic extension are characterized by parental fluids belonging to all  
489 groups (*a*, *b*, *c*, *d*, and *e*; Table 2; Fig. 8b). Cold fluids of groups *a* and *b* are associated with the infiltration of  
490 meteoric water, a common process within rocks undergoing dilation in extensional settings (Sibson, 2000;  
491 Doglioni et al., 2014; Curzi et al., 2024). Host-rock buffered fluids of group *c*, characterized by geochemical  
492 and thermal equilibrium with the host rocks, suggest near-surface fluid circulation enhanced by rapid dilation  
493 and increased porosity of rocks (Sibson 2000; Curzi et al., 2024). Hot fluids of group *d*, in geochemical  
494 equilibrium and in thermal disequilibrium with the host rocks, are consistent with the ascent of fluids from  
495 deeper structural levels during extensional faulting (Sibson 2000; Curzi et al., 2024). The proposed fluid  
496 source is the Apulian Carbonate Platform, which is separated from shallower units by the impermeable Irpinia  
497 *mélange*. The hot fluids of group *e* are interpreted as a mixing of fluid groups *a* and *d*, and therefore indicate  
498 structural connections between the Apulian Carbonate Platform and the near surface.

499 Spatial distribution of fluid groups within the Val d'Agri Basin during post-orogenic extensional tectonics (Fig.  
500 8b) highlights the key role played by the tectonic *mélange* on fluid flow. Where the *mélange* is thin, mainly  
501 below the MMFS and site 1 (Butler et al., 2004; Catalano et al., 2004), the fluid systems hosted in the units  
502 above and below the *mélange* were frequently connected by extensional-transtensional faults, which could  
503 easily breach the *mélange* (Candela et al., 2015; Schirripa Spagnolo et al., 2024). This is testified by the wide  
504 occurrence of mineralizations precipitated from fluid groups *d* and *e* (Fig. 8b). In the same areas, the

505 occurrence of mineralizations from groups *a*, *b*, and *c* testifies to a spatial and/or temporal lack of connection  
506 across the Irpinia mélange. The described fluid circulation system is consistent with the present-day natural  
507 leakage of hydrocarbons along the MMFS and in other localities of the Val d'Agri Basin (e.g., Colangelo et al.,  
508 2005; Beaubien et al., 2023).

509 Conversely, in areas characterized by thicker tectonic mélange (mainly below the EAFS; Fig. 8), the fluid  
510 systems hosted in the units above and below the mélange were generally compartmentalized, allowing the  
511 formation and preservation of the hydrocarbon trap below the Irpinia mélange (Hager et al., 2021). This  
512 process is consistent with widespread calcite precipitation from meteoric waters (groups *a* and *b*; Fig. 8b)  
513 and occasionally from host-rock buffered fluids (group *c*; Fig. 8b). It is also true, however, that at sampling  
514 site 7 (EAFS; Fig. 8b), below which the mélange thickness is > 1 km and productive oil reservoirs are located  
515 (Hager et al., 2021), mineralizations precipitated from hot fluids originated from the Apulian Carbonate  
516 Platform (fluid group *d*). This exceptional case might have been ruled by the valve behavior of major regional  
517 faults. Specifically, during strong seismic events, these faults may have allowed the hard-linkage, across the  
518 mélange, between shallow and deep structures and the related upward leakage of fluids (e.g., Sibson, 2000).  
519 The occurrence of the productive hydrocarbon reservoir below the EAFS indicates that, over the long term,  
520 extensional faulting has not substantially interrupted the continuity of the Irpinia mélange. However, we  
521 propose that (likely strong) earthquakes were able to episodically breach the mélange, allowing the ascent  
522 of fluids from the Apulian Carbonate Platform (Fig. 8b).

523 Similar occasional ascents of trapped fluids from hydrocarbon reservoirs due to seismic fault-valve action  
524 were described in the literature (e.g., Jin et al., 2008), and explained by the plastic-to-brittle transition of seal  
525 rocks with higher strain rates (e.g., Cheng & Ben-Zion, 2019). The role of earthquakes in controlling fluid  
526 upwelling from the Apulian Carbonate Platform is also emphasized by some studies suggesting that the  
527 initiation of the natural emissions of hydrocarbons near Tramutola village was related to the 1857 Mw7.1  
528 earthquake (Colangelo et al., 2005).



## 529        **6. Conclusions, implications and recommendations**

530    The paleo-fluid circulation within the Val d'Agri Basin was significantly influenced by the prevailing tectonic  
531    regime (compressive vs. extensional), the stratigraphic-structural architecture, and the occurrence of seismic  
532    events. During orogenic shortening, the progressive stacking of multiple thrust sheets substantially  
533    prevented vertical fluid circulation. Conversely, during the ongoing post-orogenic extension, the vertical  
534    circulation of fluids was predominantly controlled by the compartmentalization efficiency of the Irpinian  
535    mélange. This efficiency was improved by the increased thickness of the mélange but was compromised by  
536    faulting and associated seismic events, which promoted the upward ascent of deep fluids.

537    These results highlight the importance of considering the regional tectonic setting for hydrocarbon  
538    extraction, subsurface gas storage, or earthquake hazard mitigation. In active compressional settings, an  
539    impermeable horizon is more likely to remain intact under natural conditions, implying that fluids injecting  
540    below the impermeable horizon during extraction activities could significantly increase fluid pressure,  
541    potentially inducing earthquakes.

542    In active post-orogenic extensional settings, fluid accumulation is related to the orogenic architecture where  
543    thrust sheets may act as sealing horizons. However, active extensional faulting and associated seismic events  
544    may threaten the integrity of sealing horizons. The presence of sealing horizons and fluid traps, combined  
545    with the man-induced variations of pore fluid pressure, could also cause earthquakes, as observed in the Val  
546    d'Agri Basin (Improta et al., 2017; Hager et al., 2021). During extensional faulting, particularly in the case of  
547    strong seismic events, faults could serve as conduits for the ascent of pressurized fluids, which could be  
548    polluting or dangerous. This process will be influenced by the thickness of the sealing horizon(s).

549    In conclusion, hydrocarbon extraction and subsurface gas storage in seismically active areas present unique  
550    challenges and risks that must be addressed to reduce potential seismic or environmental hazards. Therefore,  
551    in geological settings comparable to the Val d'Agri Basin, where exploitation or storage activities are ongoing  
552    (e.g., the Pannonian Basin in Hungary, Czauer and Madl-Szonyi, 2011, the Bohai Bay Basin in China, Song et  
553    al., 2022, and the Nispiro Field in Mexico, Bourdet et al., 2010), we recommend rigorous seismic monitoring  
554    and risk assessment.

555

556 **7. Declaration of Generative AI and AI-assisted technologies in the writing process**

557 During the preparation of this work the author(s) used ChatGPT in order to improve readability and fluency.

558 After using this tool/service, the author(s) reviewed and edited the content as needed and take(s) full

559 responsibility for the content of the publication.

560

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