

# Synergistic Role of Water and Oxygen Leads to Degradation in Formamidinium-Based Halide Perovskites

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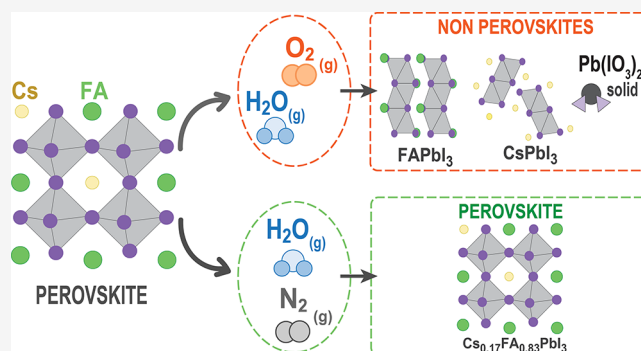


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**ABSTRACT:** Mixed-cation metal halide perovskites have shown remarkable progress in photovoltaic applications with high power conversion efficiencies. However, to achieve large-scale deployment of this technology, efficiencies must be complemented by long-term durability. The latter is limited by external factors, such as exposure to humidity and air, which lead to the rapid degradation of the perovskite materials and devices. In this work, we study the mechanisms causing Cs and formamidinium (FA)-based halide perovskite phase transformations and stabilization during moisture and air exposure. We use *in situ* X-ray scattering, X-ray photoelectron spectroscopy, and first-principles calculations to study these chemical interactions and their effects on structure. We unravel a surface reaction pathway involving the dissolution of FAI by water and iodide oxidation by oxygen, driving the Cs/FA ratio into thermodynamically unstable regions, leading to undesirable phase transformations. This work demonstrates the interplay of bulk phase transformations with surface chemical reactions, providing a detailed understanding of the degradation mechanism and strategies for designing durable and efficient perovskite materials.



## 1. INTRODUCTION

Formamidinium [CH(NH<sub>2</sub>)<sub>2</sub>, FA] metal halide perovskites have emerged as promising materials for solar cell applications due to their exceptional light-harvesting properties, producing high power conversion efficiencies (PCEs) of over 26%.<sup>1</sup> However, one of the most significant challenges is their poor stability, particularly in the presence of moisture (H<sub>2</sub>O) and oxygen (O<sub>2</sub>), which can trigger phase transformations to nonperovskite structures during solar cell fabrication and operation.<sup>2</sup> To overcome this challenge, mixed ion perovskites, including Cs<sub>x</sub>FA<sub>1-x</sub>PbI<sub>3</sub> (CsFA), have been shown to provide improved stability compared to their single-cation counterparts.<sup>2–4</sup> However, CsFA perovskites also degrade when exposed to ambient air, where the perovskite structure transforms into nonperovskite phases such as one of the hexagonal FAPbI<sub>3</sub> (2H) and orthorhombic δ-CsPbI<sub>3</sub> (δCs) structures.<sup>5,6</sup> Perovskite degradation due to water exposure<sup>7–11</sup> and photo-oxidation under illumination in oxygen is widely recognized.<sup>12–15</sup> Nonetheless, we lack a clear understanding of the mechanisms that lead to phase instabilities in these FA-rich perovskites due to water and oxygen interactions. Therefore, a fundamental understanding of the mechanisms causing perovskite phase transformations is crucial for developing durable and efficient metal halide perovskites for solar cell applications.

Herein, we investigate the origin of structural phase instability of FA-based perovskites upon exposure to H<sub>2</sub>O, with air (H<sub>2</sub>O/air) or nitrogen (H<sub>2</sub>O/N<sub>2</sub>) as carrier gases, using *in situ* grazing incidence wide-angle X-ray scattering (GIWAXS). We study the surface chemistry and propose a mechanism to explain phase transformations when the perovskites are exposed to H<sub>2</sub>O and O<sub>2</sub> by using X-ray photoelectron spectroscopy (XPS) and density functional theory (DFT) calculations. We find that the degradation rate is considerably slower when the perovskite is exposed to H<sub>2</sub>O/N<sub>2</sub> when compared to H<sub>2</sub>O/air. Our results show that a critical synergy between H<sub>2</sub>O and O<sub>2</sub> in air is needed to accelerate the undesired phase transformations in perovskites. The H<sub>2</sub>O molecules dissolve FAI on the perovskite surface, leading to volatilization of the iodide and FA<sup>+</sup> cations. The O<sub>2</sub> may then interact with the exposed lead-iodide-rich surfaces, oxidizing iodide and forming the thermodynamically favored iodate

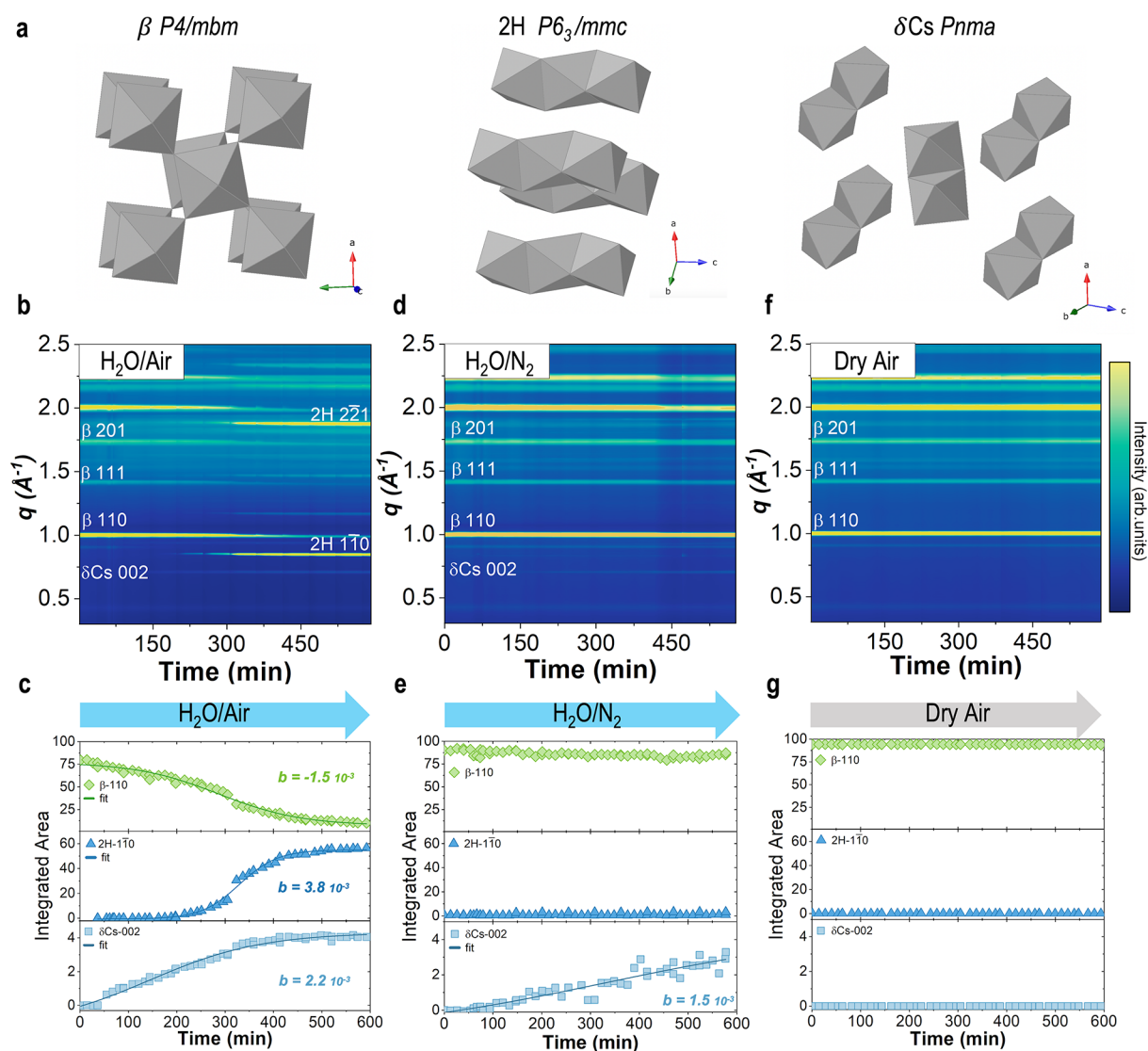
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**Figure 1.** Humidity-induced structural phase transformations measured by *in situ* GIWAXS. The structural phases analyzed are (a) tetragonal  $\beta$ -perovskite of space group  $P4/mbm$  (left), 2H  $\text{FAPbI}_3$  hexagonal nonperovskite phase of space group  $P6_3/mmc$  (center), and orthorhombic  $\delta\text{-CsPbI}_3$  nonperovskite phase of space group  $Pnma$  (right). Phase transformations over time of CsFA  $\beta$ -perovskite exposed to (b)  $\text{H}_2\text{O}/\text{air}$ , (d)  $\text{H}_2\text{O}/\text{N}_2$ , and (f) dry air, where (c, e, and g) are the corresponding integrated areas of the main scattering peaks of each phase.

species,  $\text{IO}_3^-$ , which bonds to surface Pb ions. Lead(II) iodate,  $\text{Pb}(\text{IO}_3)_2$ , can then form and segregate at the surface, leaving  $\text{PbI}_2$  vacancies behind and allowing  $\text{H}_2\text{O}$  molecules to further infiltrate the crystal structure. The continuous removal of FAI causes a local imbalance between  $\text{Cs}^+$  and  $\text{FA}^+$ , destabilizing the CsFA phase and favoring a transformation to the 2H and  $\delta\text{Cs}$  phases. The degradation proceeds both in dark and light conditions, with light accelerating the formation of secondary phases. Finally, we show that a surface treatment using a hydrophobic top layer of phenethylammonium iodide (PEAI) effectively stabilizes the perovskite phase even in the presence of  $\text{H}_2\text{O}/\text{air}$ . Solar cells made of CsFA-PEAI films exhibit stable PCEs even after exposure to  $\text{H}_2\text{O}$  and air. Our study provides structural and atomistic insights into the phase instability in FA-based perovskites when exposed to humid air conditions and provides surface passivation strategies to stabilize the perovskite phase for highly stable and efficient solar cells.

## 2. RESULTS

**2.1. Structural Phase Transformations.** To understand the structural phase transformations in FA-based perovskites, we used *in situ* GIWAXS, as shown in Figures 1 and S1, where we exposed CsFA films to a relative humidity of  $\sim 100\%$  and in dark conditions (lights off). Initially, we analyzed the structural phases without  $\text{H}_2\text{O}$  or air exposure (Le Bail refinement, Figure S2, Table S1), resulting in a mixed-cation (Cs 17%—FA 83%) tetragonal perovskite phase ( $\beta$ ) with a space group  $P4/mbm$ <sup>5,16</sup> (Figure 1a). Figure 1b shows the *in situ* GIWAXS scattering patterns as a function of time exposed to  $\text{H}_2\text{O}/\text{air}$  for 600 min. In  $\text{H}_2\text{O}/\text{air}$ , the  $\beta$  phase (mixed-cation) transformed into two different phases, namely, the single-cation  $\text{FAPbI}_3$  hexagonal phase (2H)<sup>5</sup> and  $\text{CsPbI}_3$  orthorhombic phase ( $\delta\text{Cs}$ )<sup>3</sup> (crystal structures shown in Figure 1a). The 2H and  $\delta\text{Cs}$  are both nonperovskite phases given the lack of corner-sharing octahedra. Figure 1c shows the quantified peak evolution of the integrated area of the main scattering peak of the perovskite ( $\beta$ ) and nonperovskite (2H,  $\delta\text{Cs}$ )

phases. We observe that  $\delta$ Cs forms at around 50 min of exposure to H<sub>2</sub>O/air, while 2H forms after 200 min of exposure, but at a faster rate. A kinetic model was used to quantify the phase transformation rates, shown in Figures S3 and S4, where we fit a rate constant  $b$ . The negative  $b$  for the  $\beta$ -110 peak indicates the loss of the  $\beta$ -perovskite phase in H<sub>2</sub>O/air. The positive  $b$  for the 2H and  $\delta$ Cs peaks reveals the appearance of the nonperovskite phases and a faster formation rate for 2H compared to  $\delta$ Cs.

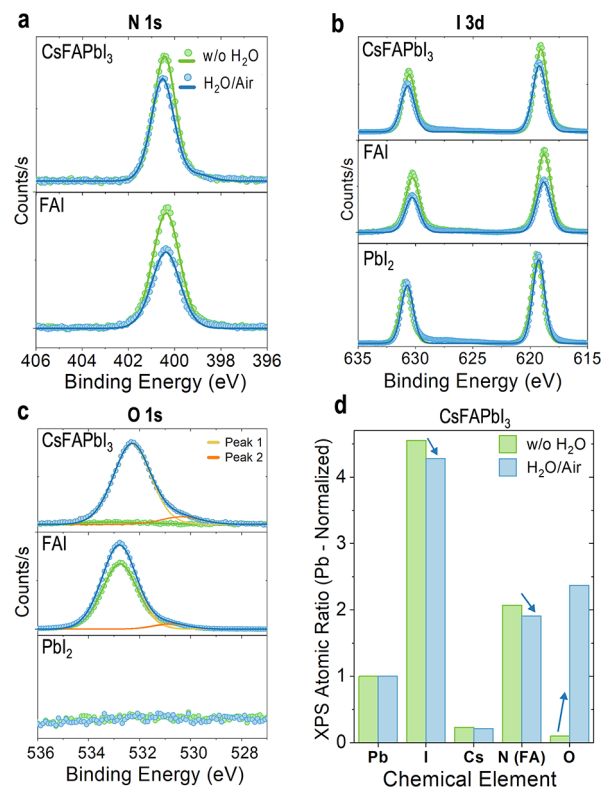
To isolate the effect of H<sub>2</sub>O from O<sub>2</sub>, we performed *in situ* GIWAXS experiments by exposing the CsFA films to H<sub>2</sub>O/N<sub>2</sub>, as shown in Figure 1d. It is important to study the role of H<sub>2</sub>O alone, considering that O<sub>2</sub> has been key in the photo-oxidation of methylammonium lead iodide (MAPbI<sub>3</sub>) under illumination.<sup>17–19</sup> Remarkably, when CsFA is exposed to H<sub>2</sub>O/N<sub>2</sub>, the  $\beta$  phase does not change, and a 2H phase does not appear. This is in contrast with the samples exposed to H<sub>2</sub>O/air, which suffered phase transformations after the same exposure time. The peak evolution in Figure 1e shows a minor decrease in the  $\beta$ -perovskite integrated area from 92 to 85 with no formation of the 2H phase. However, the  $\delta$ Cs phase still forms with a rate constant  $b$  of 1.5, compared to 2.2 in H<sub>2</sub>O/air (Figure S4), confirming a slower  $\delta$ Cs phase formation in H<sub>2</sub>O/N<sub>2</sub>.

To isolate the role of oxygen (from air) from that of the water molecule, we exposed the CsFA films to dry air only. Interestingly, from *in situ* GIWAXS (Figure 1f) and the time evolution of the main peaks (Figure 1g), the phase transformations previously seen do not occur. This reveals that the amount of photo-oxidation induced by the X-rays alone is not sufficient to induce the formation of secondary nonperovskite phases. The effects induced by humidity or dry air exposure are independent of the substrate we use (Figure S5). These results show that the CsFA phase instability is not just due to humidity exposure but is accelerated by O<sub>2</sub>. These results are further corroborated by additional analyses of the *in situ* GIWAXS experiments (Figures S5 and S6). It is worth noting that X-ray beam-induced damage due to prolonged exposure during *in situ* experiments can provide artifacts to measurements, as it has been shown in a wide variety of materials.<sup>20,21</sup> Thus, to deconvolute potential effects on the structure caused by X-rays, we measured GIWAXS in a separate isolated location of the same film (without X-ray exposure) to make sure the peak intensity did not change due to beam damage. The second isolated location was measured in parallel and with lower beam exposure, given that the data were taken with longer delays between measurements. The data for the isolated location is summarized in Figure S7 and shows a very similar trend to the consecutive measurements shown in Figure 1.

To study the structural changes at the surface of CsFA films exposed to H<sub>2</sub>O/air, we analyzed the *in situ* GIWAXS with an incident angle below the critical angle (Figure S8). We observed the same transformation from perovskite into nonperovskites on the surface as in the bulk (Figure 1). Furthermore, analyzing the formation of the nonperovskite phases 2H and  $\delta$ Cs, we calculated a larger rate constant  $b$  at the surface (Figure S8), evidence of a faster transformation into nonperovskites. In addition, areas with more charging were observed by scanning electron microscopy (Figure S9), suggesting the formation of new phases. UV–vis spectroscopy shows a larger band gap for the CsFA film after H<sub>2</sub>O/air exposure, which may be due to the absorption from the 2H phase (Figure S10).<sup>22,23</sup> The emission observed from both

photoluminescence (PL) and carrier lifetime from time-resolved PL (TRPL) is reduced, as expected with the conversion to nonperovskite phases (Figure S10).

**2.2. Surface Chemistry.** To investigate the chemical changes at the surface, we performed XPS of the CsFA films without exposure and after H<sub>2</sub>O/air exposure (Figure 2).<sup>21</sup>

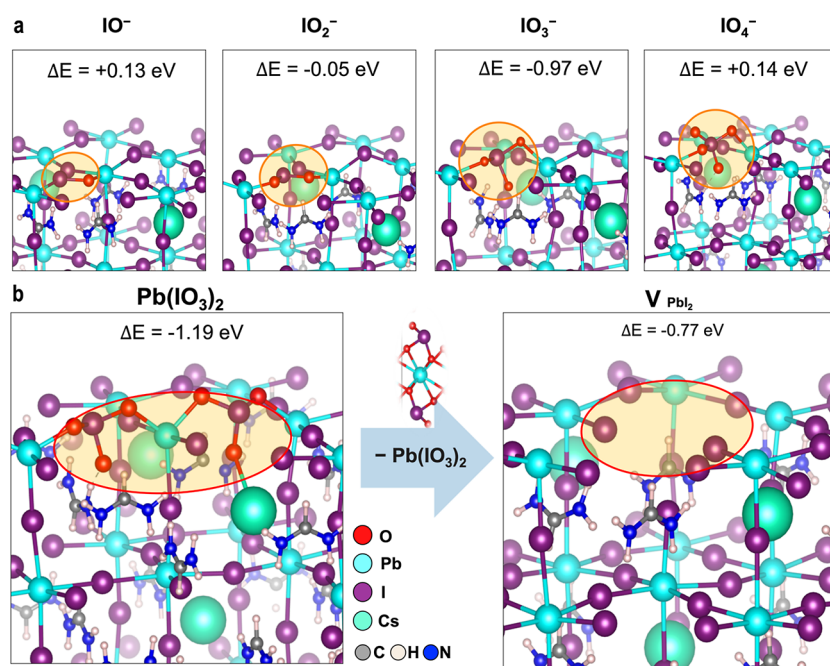


**Figure 2.** XPS spectra of the peaks: (a) N 1s, (b) I 3d, and (c) O 1s of CsFA perovskite, FAI, and PbI<sub>2</sub> films without (w/o) and after H<sub>2</sub>O/air exposure. For the CsFA perovskite films, panel (d) shows the atomic ratio of iodine(I), cesium (Cs), nitrogen (N, FA), or oxygen (O) normalized to lead (Pb), for pristine films and after H<sub>2</sub>O/air exposure. Full peak deconvolution and details for the CsFA films can be found in Figures S10 and S11 and Tables S2 and S3, for FAI films in Figure S12, Tables S4 and S5, and for PbI<sub>2</sub> films in Figure S13 and Tables S6 and S7. CsFA perovskite films w/o H<sub>2</sub>O exposure (green) were fabricated in a nitrogen glovebox and exposed to a room atmosphere while mounting the XPS measurement. CsFA films after H<sub>2</sub>O/air exposure (blue) were fabricated in a nitrogen glovebox, exposed to humidity, and then measured.

After exposure to H<sub>2</sub>O/air, the N 1s peak (Figure 2a, CsFA) decreases by 5.6% in atomic content (Tables S2 and S3), evidence of the FA<sup>+</sup> volatilization, as the N 1s peak corresponds to the C=N bond of the FA.<sup>24</sup> The I 3d peak from iodine decreases by 11.4% in atomic content after exposure to H<sub>2</sub>O/air (Figure 2b, CsFA), suggesting the loss of FAI at the surface and the additional loss of iodine from elsewhere in the structure.<sup>19,25,26</sup>

The exposure of CsFA films to H<sub>2</sub>O/air influences the oxygen signal. An increase in the intensity of the peak signal of the O 1s is observed after exposure (Figure 2c), corresponding to an 18% atomic content (Table S3). The O 1s peak (peak 1) is centered at 532.3 eV, suggesting the existence of adsorbed O<sub>2</sub> molecules.<sup>27</sup> Peak 1 may be attributed to the formation of hydroxides (OH<sup>-</sup>) or carbonates (CO<sub>3</sub><sup>2-</sup>) expected from the exposure to H<sub>2</sub>O/air.<sup>28</sup> The C 1s peaks also suggest carbon–





**Figure 3.** DFT calculations of iodide oxidation and superoxide formation on the perovskite surface. Panel (a) visualizes the formation of  $\text{IO}_n^-$  species upon oxidation of surface iodide ions [Rx. 1]; from left to right: hypiodite,  $\text{IO}^-$ ; iodite,  $\text{IO}_2^-$ ; iodate,  $\text{IO}_3^-$ ; and periodate,  $\text{IO}_4^-$ . All oxidized species are explicitly highlighted, and reaction energies are given (see the computational methods for details). (b) Reaction mechanism [Rx. 2 and 3] of (left) lead(II) iodate,  $\text{Pb}(\text{IO}_3)_2$ , formation by consumption of oxygen molecules and (right) removal of  $\text{Pb}(\text{IO}_3)_2$  resulting in a surface vacancy  $\text{V}_{\text{PbI}_2}$ . The following color code is used for the atomic representations: purple, I; cyan, Pb; blue, N; green, Cs; gray, C; white, H; and red, O.

oxygen complexes such as carbonates after  $\text{H}_2\text{O}$ /air exposure (Figure S11).<sup>27</sup> The peak assigned to C–C or C–H bonds increases after  $\text{H}_2\text{O}$ /air exposure, possibly from adventitious carbon (Figure S12). Figure 2c also shows a small peak 2, centered at 530.4 eV, suggesting the formation of Pb-oxides.<sup>27</sup> Figure 2d summarizes the changes in the atomic ratio of iodine, cesium, nitrogen, and oxygen compared to Pb on CsFA films after  $\text{H}_2\text{O}$ /air exposure, showing the volatilization of nitrogen and iodine, and an increase in oxygen content.

To further understand the chemical surface species after  $\text{H}_2\text{O}$ /air exposure, we deposited FAI and  $\text{PbI}_2$  thin films and characterized their surface chemistries by XPS (Figure 2a–c FAI,  $\text{PbI}_2$ , and Figures S13 and S14). After  $\text{H}_2\text{O}$ /air exposure, the FAI films show a decrease in the atomic content from N 1s and I 3d, confirming the volatilization of FAI (Figure 2a,b, FAI, and Tables S4 and S5). An increase in the intensity of the O 1s peak is observed for the FAI films (Figure 2c, FAI). The FAI films without  $\text{H}_2\text{O}$  exposure show an O 1s peak centered at 532.8 eV in Figure 2c (FAI), suggesting that atmospheric  $\text{H}_2\text{O}$  and  $\text{O}_2$  molecules are adsorbed to the FAI films when the samples are mounted in the XPS, confirming the high hydrophilicity of FAI. The deconvolution of the O 1s peak shows a second peak (peak 2) centered at 530.8 eV, indicating that this peak may result from FA exposure to ambient conditions. Therefore, the O 1s peak 2 in Figure 2c (CsFA) can be attributed to either Pb-oxides and to I- or FA-based oxides. In contrast, the  $\text{PbI}_2$  films before and after  $\text{H}_2\text{O}$ /air exposure do not show changes in Pb 4f (Figure S14) or I 3d (Figure 2b,  $\text{PbI}_2$ ), and there is no oxygen from the O 1s spectrum (Figure 2c,  $\text{PbI}_2$  and Tables S6 and S7). We studied the chemistry changes in the bulk by X-ray fluorescence (XRF) elemental mapping, which also shows the loss of iodine after  $\text{H}_2\text{O}$ /air exposure (Figure S15), agreeing with the XPS results

(Figure 2b). Fourier transform infrared spectra (FTIR) showed no oxygen in the vibrational modes of CsFA powders and films (Figures S16 and S17, Table S8). Thus, we are confident that the  $\text{H}_2\text{O}$ /air-induced CsFA phase transformations begin at the surface through the interaction of oxygen with the perovskites.

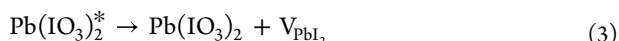
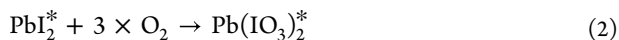
**2.3. Chemical Reaction Mechanism.** Previous studies have suggested adverse effects of  $\text{H}_2\text{O}$  on lead-iodide perovskite surfaces leading to the degradation of the material.<sup>6,17,29,30</sup> Molecular dynamics simulations of  $\text{MAPbI}_3/\text{H}_2\text{O}$  interfaces have shown the fast dissolution of MAI-terminated surfaces, as  $\text{H}_2\text{O}$  molecules break the bond between surface  $\text{I}^-$  and underlying  $\text{Pb}^{2+}$  ion resulting in the removal of  $\text{I}^-$ , accompanied by the desorption of the  $\text{MA}^+$ .<sup>31,32</sup> Despite FA being less polar than MA,<sup>33</sup> the same mechanism is expected to occur on FAI-terminated surfaces. Previous results showed the dissolution of  $\text{FAPbI}_3$  with water vapor,<sup>34</sup> explaining the release of  $\text{I}^-$  and  $\text{FA}^+$  from the CsFA perovskite surface<sup>31</sup> and within the FAI films (Figure 2a,b,d) after exposure to  $\text{H}_2\text{O}$ /air.<sup>32</sup> The removal of CsI or FAI leads to a  $\text{PbI}_2$ -terminated surface that is more resistant to degradation by water alone.<sup>31</sup>

To understand the phase transformation mechanisms, we performed DFT calculations on the interactions of oxygen molecules with the  $\text{PbI}_2$ -terminated CsFA perovskite surface (Figure S18). Our calculations suggest that the  $\text{O}_2$  adsorption is favored at the  $\text{PbI}_2$ -terminated surface ( $E_{\text{ads}} = -0.03$  to  $-0.22$  eV, Table S9) compared to a CsFAI-terminated surface ( $E_{\text{ads}} = 0.06$  to  $0.36$  eV, Table S10), while  $\text{H}_2\text{O}$  may further support  $\text{O}_2$  adsorption. Note that the  $\text{PbI}_2$ -terminated surface is made of undercoordinated  $\text{Pb}^{2+}$ , which makes it more hydrophilic than the  $\text{PbI}_2$  crystal phase (Figure 2c) made of fully coordinated  $\text{Pb}^{2+}$ .<sup>35</sup> The direct formation of PbO units

from the adsorption of  $O_2$  appears thermodynamically unfavorable, with reaction energies of +1.09 eV (Figure S19). Therefore, we consider the oxidation of iodide ions on the surface by  $O_2$  as the starting point, as modeled by reaction 1. Here, surface iodide changes its oxidation state from  $-1$  to positive values ( $-1 + 2 \times n$ ) while oxygen atoms are in their stable  $-2$  oxidation state. We observe that, as previously suggested,<sup>36</sup> oxygen breaks the Pb–I bonds at the surface, forming a Pb–O–I<sub>ox</sub>–I<sup>−</sup> bond network, where I<sub>ox</sub> is the oxidized iodine and I<sup>−</sup> is a surface iodide in its negative charge state as shown in Figure 3a.



Our calculations show positive reaction energies for the formation of hypoiodite ( $IO^-$ ) and periodate ( $IO_4^-$ ) of 0.13 and 0.14 eV, respectively, while iodite ( $IO_2^-$ ) is slightly favored by  $-0.05$  eV. Notably, iodate ( $IO_3^-$ ) with iodide in its +5-oxidation state is strongly thermodynamically favored with a formation energy of  $-0.97$  eV, indicating an irreversible surface transformation. From  $IO_3^-$  in reaction 2, we consider the formation of lead(II) iodate,  $Pb(IO_3)_2$ , at the surface (indicated by an asterisk) by oxidizing two adjacent iodide ions sharing bonds with the same Pb surface ion. The formation energy of surface  $Pb(IO_3)_2$  is  $\Delta E = -1.19$  eV (Figure 3b), which suggests a thermodynamically favorable replacement of a surface  $PbI_2$  unit. Finally, in reaction 3, we compute the removal of surface  $Pb(IO_3)_2$  leaving a  $PbI_2$  vacancy (Figure 3b).<sup>37</sup> We predict an energy of  $-0.77$  eV for reaction 3, significantly more favorable than the direct removal of a  $PbI_2$  unit from a nonoxidized surface ( $\Delta E = 0.07$  eV, Figure S20).

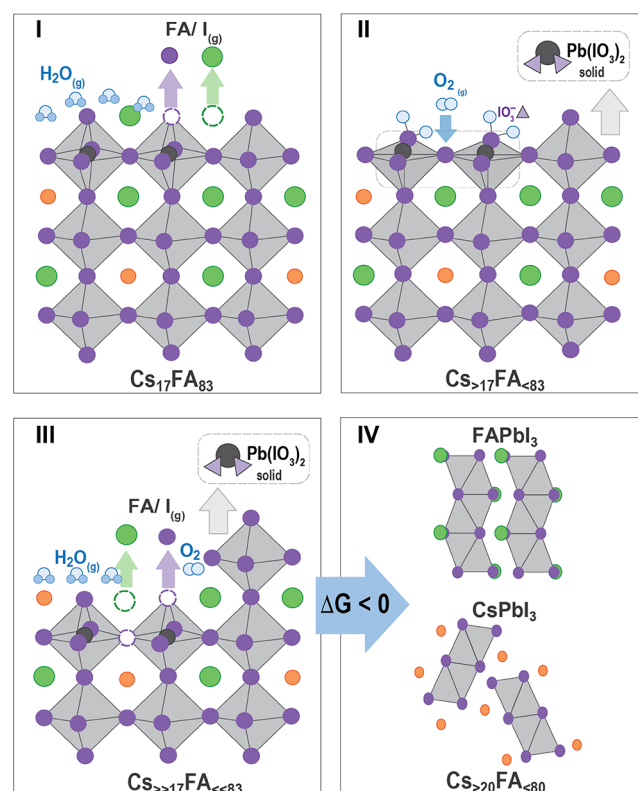


We note that the interaction of  $O_2$  with water molecules lowers the  $\pi^*$  orbitals of  $O_2$  acting as accepting orbitals in oxidation reactions (Table S10), while hydration of perovskite surfaces raises the energies of the iodide-based valence band edge (Figure S21), which results in easier oxidation of surface iodide. These results suggest that  $O_2$  can modify  $PbI_2$ -terminated surfaces by oxidizing iodide to iodate species and creating  $PbI_2$  vacancies. These vacancies may allow  $H_2O$  molecules to enter the structure and dissolve the next FAI layers in an iterative process.

Our experimental and theoretical analyses suggest that first the  $H_2O$ /air atmosphere dissolves the surface of CsFA perovskite and causes a loss in FA at the surface, likely increasing the relative amount of Cs relative to FA. Previous studies have shown that the CsFA perovskite phase becomes thermodynamically unstable when exceeding a Cs fraction of 0.2 (FA below 0.8).<sup>3,4</sup> We note that the studied composition of  $Cs_{0.17}FA_{0.83}$  is on the upper limit of Cs-molar content to form a single-phase perovskite.<sup>4,5</sup> The loss of FAI may easily shift the composition into the thermodynamically unstable phase by increasing the ratio between Cs and FA, favoring the segregation into nonperovskite phases. Iodide vacancies created by  $H_2O$  may further accelerate the phase transitions.<sup>10</sup> Light can speed up these reactions, but degradation can also happen in the dark. This is confirmed both by the trend in the XPS spectra, where exposure to  $H_2O$ /air is performed in the dark, and by GIWAXS measurements performed on the same

sample but in different spots and using different measurement frequencies (different X-ray doses), Figure S7. Previous studies have highlighted the role of superoxide on perovskite degradation under light irradiation.<sup>13,19,25,38</sup> Both room light and X-ray exposure can photoexcite charge carriers during the *in situ* GIWAXS measurements, but they only lead to a negligible amount of free charge carriers and superoxide formation (discussion in S2.5, Figures S22–S26).

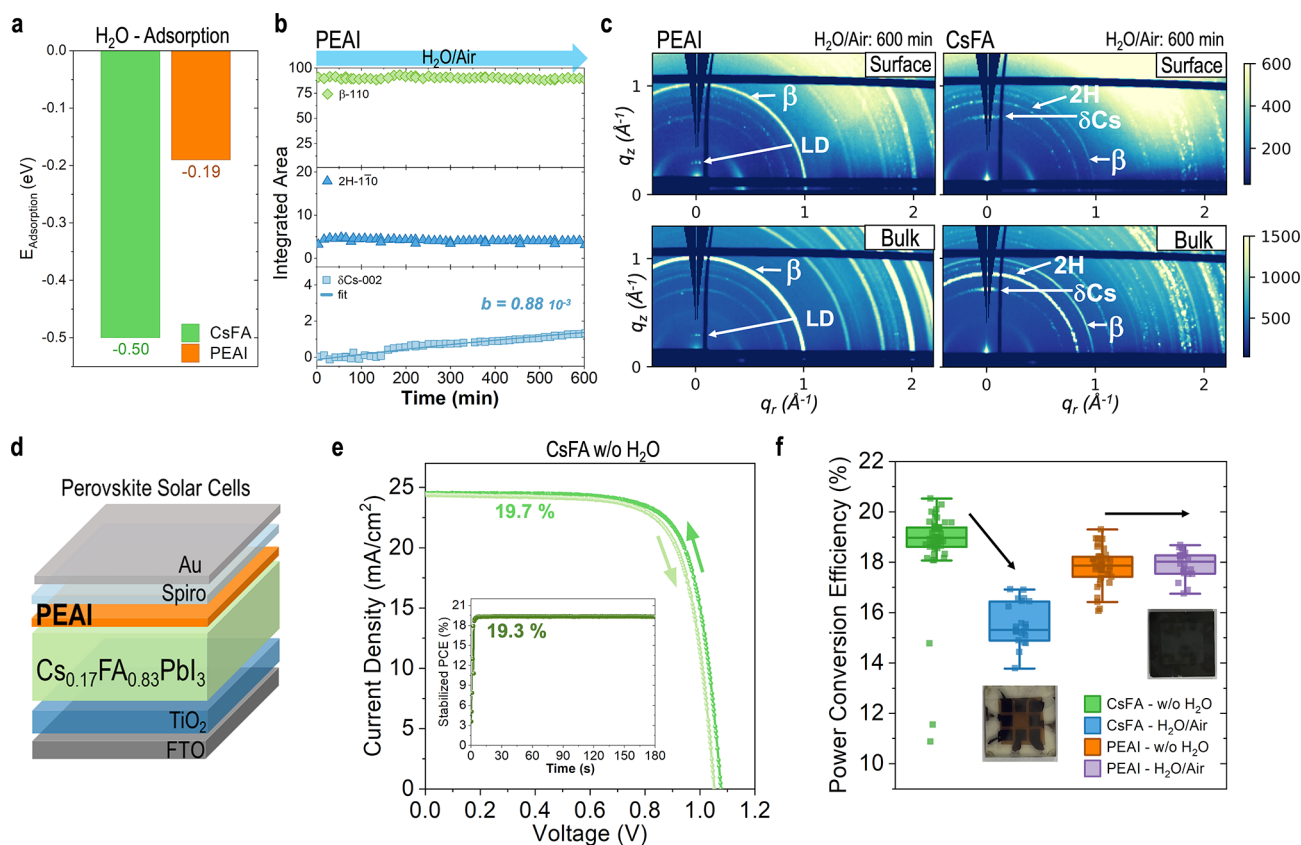
A proposed mechanism for the role of chemistry in phase transformations during exposure to  $H_2O$ /air is shown in Figure 4. First (I), water molecules dissolve FAI at FAI-rich perovskite



**Figure 4.** Proposed surface and bulk mechanism for CsFA perovskites exposed to  $H_2O$ /air. (I)  $H_2O$  is adsorbed on the CsFA surface, promoting the loss of FAI (g). Surface vacancies are created, leading to preferential oxygen binding sites, favoring the oxidation of iodide and energetically favorable formation of  $Pb(IO_3)_2$  (II), which will create a  $PbI_2$  vacancy (III). Surface vacancies and the loss of FAI will lead to faster phase segregation and phase transformations from mixed-cation perovskite into single-cation nonperovskite phases (IV).

surfaces, creating  $PbI_2$ -rich surface regions. Second (II), the  $O_2$  molecules are adsorbed and oxidize surface iodide, resulting in iodate species at the surface. Third (III), lead iodate can subsequently segregate at the surface, leaving  $PbI_2$  vacancies, which act as hotspots for further FAI dissolution by  $H_2O$ . Fourth (IV), the loss of FA changes the Cs/FA molar ratio beyond the thermodynamically stable region, forming single-cation nonperovskite phases. We emphasize that the structural phase transformations in  $H_2O$ /air are initiated at the surfaces. Thus, chemical surface treatments, such as reducing under-coordinated Pb ions at the surface, should reduce the interaction with  $H_2O$  and  $O_2$ , likely improving phase stabilization.

**2.4. Stabilizing the Perovskite Phase.** To understand the role of surface blockers, capping layers, or surface



**Figure 5.** The PEAI capping layer is used to stabilize the perovskite in H<sub>2</sub>O/air. Panel (a) shows H<sub>2</sub>O adsorption energy on a CsFA surface and CsFA-PEAI surface calculated by DFT. From GIWAXS measurements in Figure S27, (b) is the integrated area of the main scattering peak of each phase for films exposed to H<sub>2</sub>O/air. (c) Shows the 2D GIWAXS patterns from the surface and bulk measurements after 600 min of exposure to H<sub>2</sub>O/air for the (left) PEAI-treated CsFA films and (right) untreated CsFA. A PEAI-treated CsFA perovskite layer in a solar cell with (d) n-i-p architecture. Panel (e) shows the current density–voltage curve and stabilized PCE of a high-efficiency device, and (f) shows the statistics of the PCE in box plots for the CsFA-untreated and PEAI-treated, w/o and with H<sub>2</sub>O exposure, under 1 sun illumination. The inset pictures show the device after H<sub>2</sub>O/air exposure before depositing the Spiro-OMeTAD and Au layers.

passivators on phase transformations at the interface,<sup>39,40</sup> we evaluated the effect of PEAI spin-coated on a perovskite film (PEAI-treated).<sup>41,42</sup> DFT calculations show a substantial reduction in H<sub>2</sub>O adsorption energies at the aromatic PEA<sup>+</sup> (Figure 5a), resulting in hydrophobic protection that can reduce the degradation of CsFA perovskite surfaces (Figure S27). We analyzed the structural phase transformations by *in situ* GIWAXS in H<sub>2</sub>O/air (Figures S28–S30). Figure 5b shows the peak evolution of the main perovskite and nonperovskite phases (Figure 1a) as a function of exposure time to H<sub>2</sub>O/air. The main 110  $\beta$ -perovskite peak does not change in H<sub>2</sub>O/air after 600 min of exposure, suggesting that the PEAI layer prevents H<sub>2</sub>O and O<sub>2</sub> interactions at the surface compared to the untreated films in H<sub>2</sub>O/air (Figure 1b,c). Films treated with PEAI show a slight presence of 2H phase (Figure 5b), which decreases as a function of time (Figure S30). The  $\delta$ Cs phase forms when exposed to H<sub>2</sub>O/air but three times slower than the untreated films (Figure 1c) when comparing the rate constant  $b$  (Figures S2 and S30). Figure 5c shows the 2D GIWAXS after final exposure to H<sub>2</sub>O/air. For the PEAI-treated films, the  $\beta$ -perovskite 110 Debye–Scherrer ring has a high intensity at the bulk and surface. We also observe rings that we assigned to low-dimensional (LD) PEA phases, as expected.<sup>43</sup> The pristine CsFA films in Figure 5c show that the primary phases are 2H and  $\delta$ Cs.

We evaluated the effect of the PEAI treatment on the performance and stability of solar cells of the n-i-p architecture (Figure 5d). CsFA-untreated devices showed an initial PCE of 19.7% (Figure 5e). The average PCE for the CsFA-untreated solar cells dropped 3.7% after H<sub>2</sub>O/air exposure (CsFA-H<sub>2</sub>O/air), while the average PCE for the PEAI-treated devices did not decrease (PEAI-H<sub>2</sub>O/air) (Figure 5f). A change in film color was visible after the pristine films were exposed to H<sub>2</sub>O/air (images in Figure 5f). The drop of PCE in CsFA-H<sub>2</sub>O/air is mainly due to the decrease in short circuit current density (Figure S31). We attribute this decrease to the formation of wide band gap nonperovskite phases such as 2H and  $\delta$ Cs, which absorb less photons and thus transport fewer charge carriers in the solar cell. From correlative XRF and X-ray beam-induced current (XBIC) maps in Figure S32, we observe the formation of Cs-rich clusters that could correspond to  $\delta$ -CsPbI<sub>3</sub>. The  $\delta$ Cs clusters in XRF are correlated with reduced XBIC currents, in agreement with other studies.<sup>44</sup> We assessed the long-term stability under operating conditions of solar cells<sup>14,38</sup> under one sun illumination in dry air (Figures S33 and S34). Our results show an 85% decrease after 13 h in the untreated film solar cells. The PEAI-treated solar cells decreased only 30% from their initial PCE after 25 h, showing a slower degradation. This is in agreement with previous studies that highlighted the relevance of photo-oxidation of perovskite under 1 sun illumination and O<sub>2</sub> exposure and



further suggests increased robustness to the exposure to H<sub>2</sub>O/air of PEAI-treated surfaces (Figure S32).<sup>13,19,25,38</sup> In contrast, the solar cells exposed to dry nitrogen showed little to no degradation (Figure S33), suggesting that oxygen in air is key to the degradation of the solar cells.

### 3. CONCLUSIONS

Exposing mixed-cation CsFA perovskites to H<sub>2</sub>O/air leads to undesired structural phase transformations, unlike the slower degradation observed in H<sub>2</sub>O/N<sub>2</sub> and little to no degradation in dry air and the dark. When exposed to H<sub>2</sub>O/N<sub>2</sub>, the CsFA perovskite degrades more slowly by dissolving FAI molecules from the surface. However, in the presence of both H<sub>2</sub>O and air, after the volatilization of the FAI molecules, the O<sub>2</sub> from air oxidizes surface iodide ions, forming lead(II) iodate. This further causes the formation of PbI<sub>2</sub> vacancies, which act as hotspots for H<sub>2</sub>O to enter the structure and that lead to the volatilization of additional FAI molecules. This results in the loss of FAI molecules in an iterative process. This alteration in the local composition leads the Cs/FA ratio beyond the energetically stable region, producing a thermodynamic force that drives the phase transformation from perovskite to nonperovskite phases. This phase transformation process starts at the surface, where H<sub>2</sub>O and O<sub>2</sub> react with surface ions. These insights into the surface chemistry and reaction mechanisms provide a foundation for designing durable and efficient solar cell materials. As a demonstration, a hydrophobic PEAI layer can be used to protect the surface from water and oxygen molecules, which prevents structural phase transformations and helps to preserve the solar cell performance.

### ■ ASSOCIATED CONTENT

#### SI Supporting Information

The Supporting Information is available free of charge at <https://pubs.acs.org/doi/10.1021/jacs.3c05657>.

All experimental methods and additional supporting results, *in situ* relative humidity GIWAXS measurements, Le Bail analysis of powder and thin-film diffraction data, lattice parameters from Le Bail analysis, integrated circular average from *in situ* relative humidity for different substrates and exposure conditions, GIWAXS measurement in a separate isolated location measured in parallel and with lower beam exposure, structural phase analysis at the surface, morphology from SEM in CsFA thin films w/o H<sub>2</sub>O exposure and after exposed ex-situ to H<sub>2</sub>O/air, optical properties of the CsFA exposed ex-situ to H<sub>2</sub>O/air, XPS spectra of CsFA films w/o exposure to H<sub>2</sub>O and exposed ex-situ to H<sub>2</sub>O/air, XPS convoluted peaks' details and atomic percent for CsFA and FAI and PbI<sub>2</sub> films w/o exposure to H<sub>2</sub>O and with exposure to H<sub>2</sub>O/air, XPS of deposited FAI and PbI<sub>2</sub> thin films, XRF maps of the molar ratios between Cs, Pb, and I in CsFA thin films, FTIR spectra of CsFA thin films and powders, FTIR peaks and assignment for powders and thin films, surface models of CsFA with PbI<sub>2</sub> and CsFAI termination, adsorption energy of O<sub>2</sub> on PbI<sub>2</sub>- and CsFAI-terminated perovskite surfaces, DFT calculation of PbO formation at a defective PbI<sub>2</sub>-terminated surface and of PbI<sub>2</sub> vacancy formation, energy of  $\pi^*$  orbitals of molecular oxygen, DOS of the perovskite slab for the pristine and hydrated surface, top and side view of the FAI-terminated perovskite and the

PbI<sub>2</sub>-terminated perovskite surface (DFT), DOS from spin-polarized DFT calculations for the pristine, hydrated, and defective surfaces after O<sub>2</sub> adsorption, superoxide formation energies, H<sub>2</sub>O adsorption on the PEAI-terminated surface of a two-dimensional PEAI<sub>2</sub>PbI<sub>4</sub> perovskite and PbI<sub>2</sub>-terminated surface of the considered CsFA perovskite, *in situ* GIWAXS in H<sub>2</sub>O/air of PEAI-treated samples, simulated diffraction files for LD PEAI cation Ruddlesden Popper phases, current-voltage curves for solar cells for the untreated CsFA perovskites and for the PEAI-treated without H<sub>2</sub>O exposure and after ex-situ H<sub>2</sub>O/air exposure, XRF with correlative XBIC done on complete CsFA solar cell devices, and long-term stability of the solar cells of CsFA without and with a top layer of PEAI (PDF)

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## Notes

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